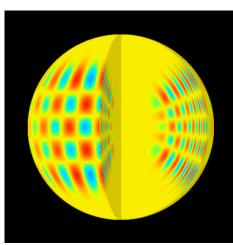
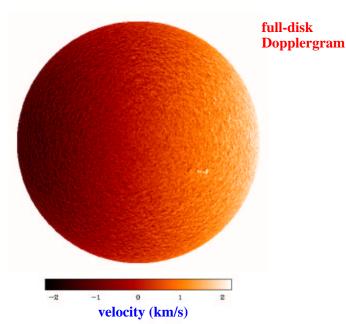
#### **HELIOSEISMOLOGY (I)**

acoustic mode in the Sun (p mode n=14, 1 -20)





## STRUCTURE OF THE SUN

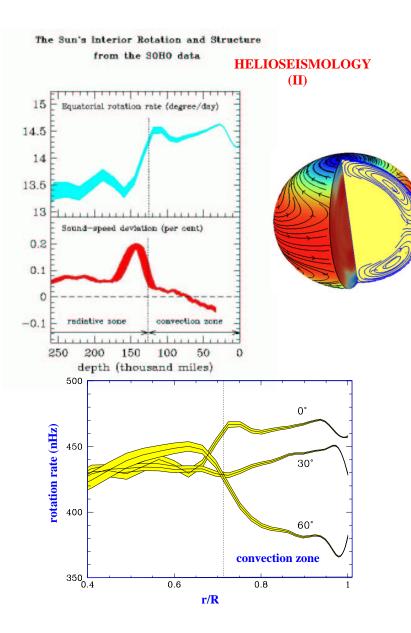
- The Sun is the only star for which we can measure internal properties  $\rightarrow$  test of stellar structure theory
- Composition (heavy elements) from meteorites
- Density, internal rotation from helioseismology
- Central conditions from neutrinos

### HELIOSEISMOLOGY

- The Sun acts as a resonant cavity, oscillating in millions of (acoustic, gravity) modes (like a bell)
- $\rightarrow$  can be used to reconstruct the internal density structure (like earthquakes on Earth)
- oscillation modes are excited by convective eddies
- periods of typical modes: 1.5 min to 20 min
- $\bullet$  velocity amplitudes:  $\sim 0.1\,m/s$
- need to measure Doppler shifts in spectral lines relative to their width to an accuracy of  $1:10^6$ 
  - > possible with good spectrometers and long integration times (to average out noise)

## Results

- density structure, sound speed
- $\bullet$  depth of outer convective zone:  $\sim 0.28\,R_\odot$
- rotation in the core is slow (almost like a solid-body)
   → the core must have been spun-down with the envelope (efficient core–envelope coupling)



## SOLAR NEUTRINOS

- Neutrinos, generated in solar core, escape from the Sun and carry away 2-6% of the energy released in H-burning reactions
- they can be observed in underground experiments
   → direct probe of the solar core
- neutrino-emitting reactions (in the pp chains)

 $\begin{array}{rrrr} {}^{1}\!H + \!\!\!^{1}\,H & \to \ {}^{2}\!D + e^{+} + \nu & E_{\nu}^{max} = 0.42\,Mev \\ {}^{7}\!Be + e^{-} & \to \ {}^{7}\!Li + \nu & E_{\nu}^{max} = 0.86\,Mev \\ {}^{8}\!B & \to \ {}^{8}\!Be + e^{+} + \nu & E_{\nu}^{max} = 14.0\,Mev \end{array}$ 

• The Davis experiment (starting around 1970) has shown that the neutrino flux is about a factor of 3 lower than predicted  $\rightarrow$  the solar neutrino problem

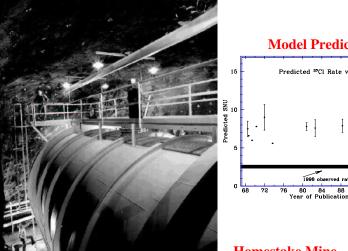
# The Homestake experiment (Davis)

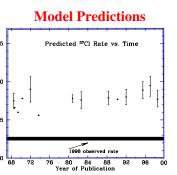
- neutrino detector: underground tank filled with 600 tons of Chlorine ( $C_2 Cl_4$ : dry-cleaning fluid)
- some neutrinos react with Cl

 $u_{\mathrm{e}} + ~^{37}\!\mathrm{Cl} 
ightarrow ~^{37}\!\mathrm{Ar} + \mathrm{e^-} - 0.81\,\mathrm{Mev}$ 

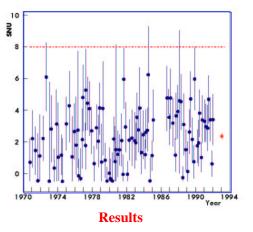
- $\bullet$  rate of absorption  $\sim 3 \times 10^{-35} \, {\rm s}^{-1}$  per  $^{37}\!Cl$  atom
- every 2 months each <sup>37</sup>Ar atom is filtered out of the tank (expected number: 54; observed number: 17)
- caveats
  - b difficult experiment, only a tiny number of the neutrinos can be detected
  - b the experiment is only sensitive to the most energetic neutrinos in the <sup>8</sup>B reaction (only minor reaction in the Sun)

#### **The Davis Neutrino Experiment**



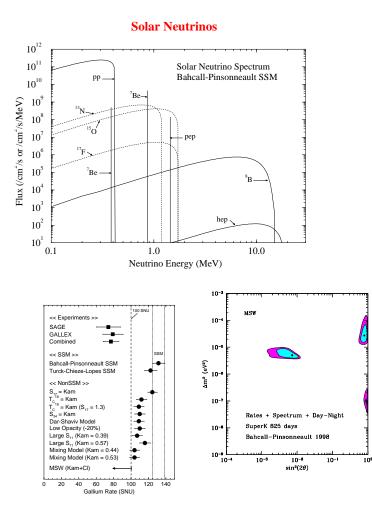


**Homestake Mine** (with Cl tank)



### **Proposed Solutions to the Solar Neutrino Problem**

- dozens of solutions have been proposed
- 1) Astrophysical solutions
  - ▷ require a reduction in central temperature of about 5 % (standard model:  $15.6 \times 10^6 \,\mathrm{K}$ )
  - ▷ can be achieved if the solar core is mixed (due to convection, rotational mixing, etc.)
  - ▷ if there are no nuclear reactions in the centre (inert core: e.g. central black hole, iron core, degenerate core)
  - b if there are additional energy transport mechanisms (e.g. by WIMPS = weakly interacting particles)
  - ▷ most of these astrophysical solutions also change the density structure in the Sun  $\rightarrow$  can now be ruled out by helioseismology
- 2) Nuclear physics
  - ▷ errors in nuclear cross sections (cross sections sometimes need to be revised by factors up to  $\sim 100)$
  - **b** improved experiments have confirmed the nuclear cross sections for the key nuclear reactions



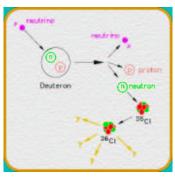
# 3) Particle physics

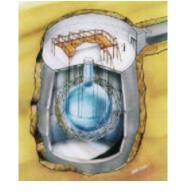
- > all neutrinos generated in the Sun are electron neutrinos
- $\triangleright$  if neutrinos have a small mass (actually mass differences), neutrinos may change type on their path between the centre of the Sun and Earth: neutrino oscillations, i.e. change from electron neutrino to  $\mu$  or  $\tau$  neutrinos, and then cannot be detected by the Davis experiment
- ▷ vacuum oscillations: occur in vacuum
- b matter oscillations (MSW [Mikheyev-Smirnov--Wolfenstein] effect): occur only in matter (i.e. as neutrinos pass through the Sun)

## **RECENT EXPERIMENTS**

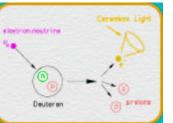
- resolution of the neutrino puzzle requires more sensitive detectors that can also detect neutrinos from the main pp-reaction
- 1) The Kamiokande experiment (also super-Kamiokande)
  - b uses 3000 tons of ultra-pure water (680 tons active medium) for
    - $u + e^- \rightarrow v + e^- \text{ (inelastic scattering)}$
  - $\triangleright$  about six times more likely for  $\nu_{\rm e}$  than  $\nu_{\mu}$  and  $\nu_{\tau}$
  - > observed flux: half the predicted flux (energy dependence of neutrino interactions?)

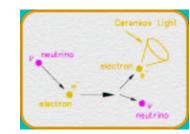
The Sudbury Neutrino Observatory

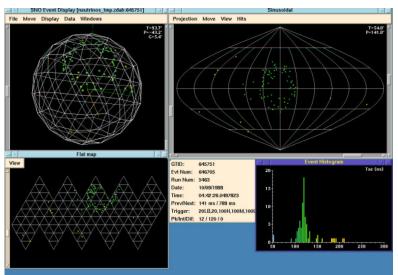




1000 tons of heavy water







- 2) The Gallium experiments (GALLEX, SAGE)
  - b uses Gallium to measure low-energy pp neutrinos directly

 $\nu_{e}+^{71}Ga\rightarrow^{71}Ge+e^{-}-0.23\,Mev$ 

- hightarrow results: about 80 ± 10 SNU vs. predicted  $132 \pm 7$  SNU (1 SNU: 10<sup>-36</sup> interactions per target atom/s)
- 3) The Sudbury Neutrino Observatory (SNO)
  - $\triangleright$  located in a deep mine (2070 m underground)
  - $\triangleright$  1000 tons of pure, heavy water (D<sub>2</sub>O)
  - $\triangleright$  in a crylic plastic vessel with 9456 light sensors/photo-multiplier tubes
  - b detect Cerenkov radiation of electrons and photons from weak interactions and neutrino-electron scattering
  - results (June 2001): confirmation of neutrino oscillations (MSW effect)?