

# The Search for the Shortest Period Binary WDs: The Latest Results from the ELM Survey



Alexandros Gianninas

sdOB7

University of Oxford

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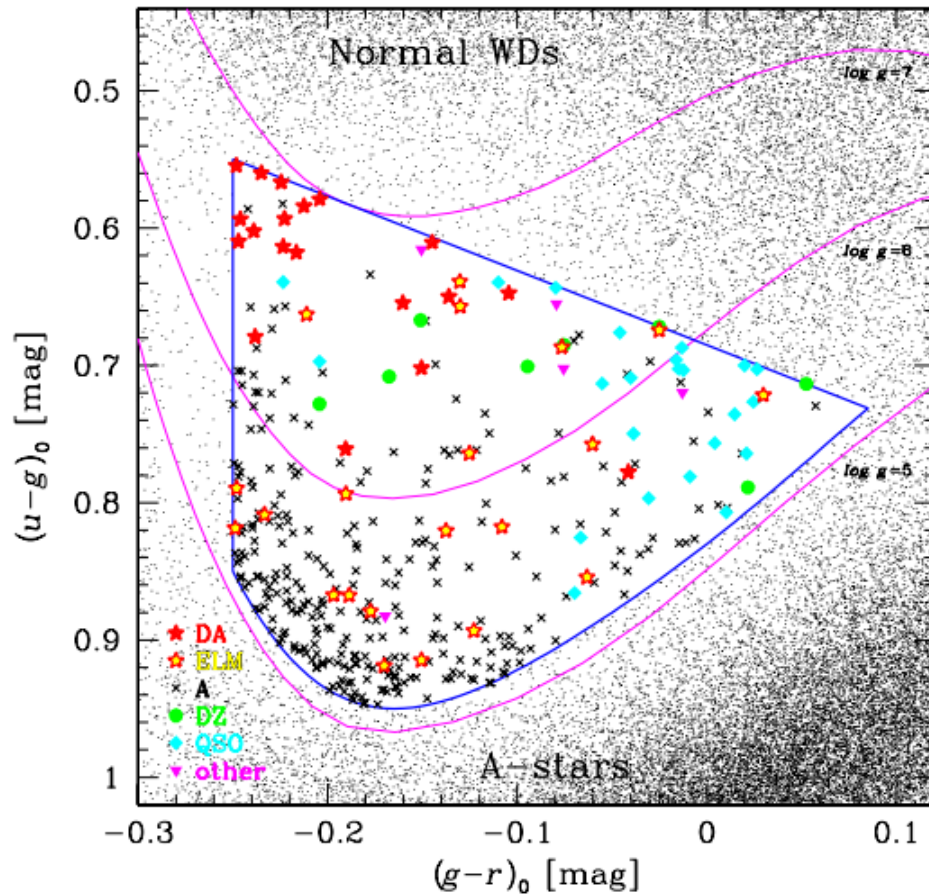
Collaborators:

M. Kilic, P. Canton, S. Barber, C. Wood (U. Oklahoma); W. R. Brown, S.J. Kenyon (SAO/CfA);  
J.J. Hermes (U. Warwick); P. Dufour, P. Bergeron (U. de Montréal)

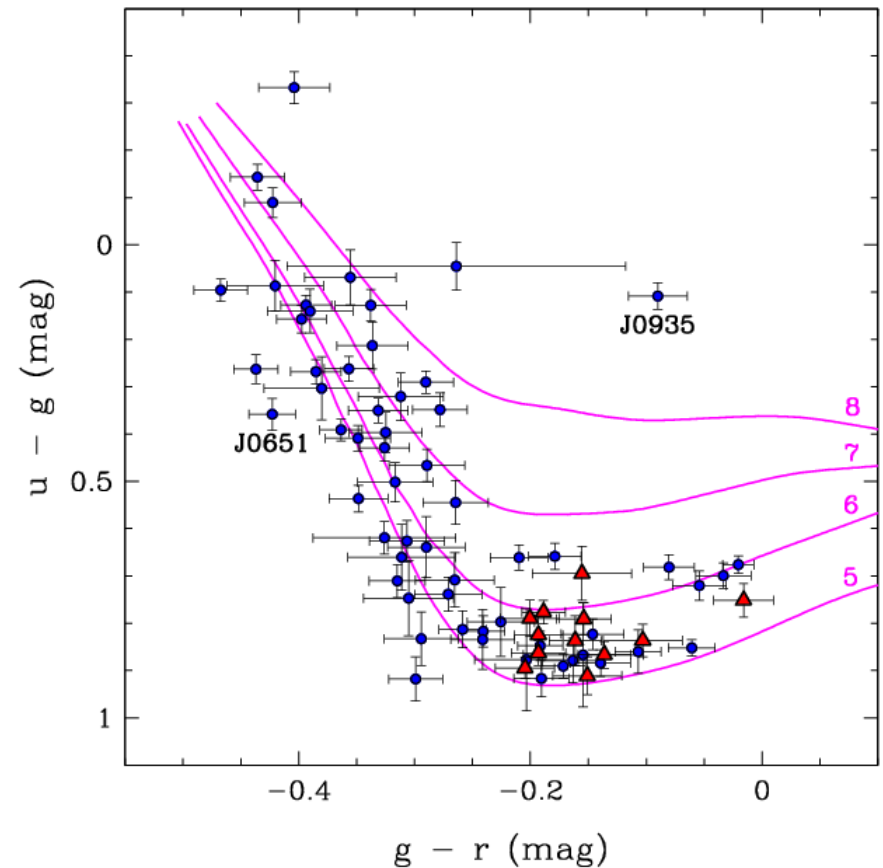
# The ELM Survey

- The ELM Survey is an ongoing, targeted search for ELM WDs
  - $M < 0.30 M_{\odot}$  ( $5 < \log g < 7$ )
  - Found in short-period ( $P < 1$  day) binary systems
- Motivation
  - Progenitors of SN Ia, .Ia, AM CVn, R CrB, **single sdBs**
  - Neutron star companions
  - Gravitational wave verification sources
  - Laboratories to test GR, tidal effects (J0651)
- Papers I – V + VI : 62 + 12 new ELM WDs

# SDSS colors work well for choosing ELM WD candidates

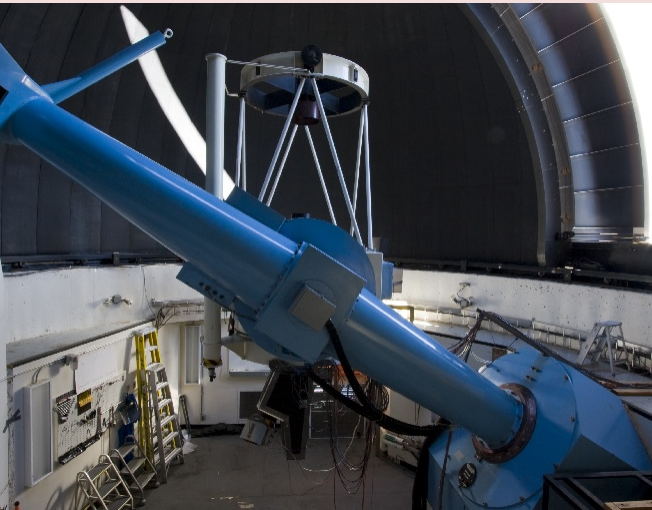


Brown et al. (2012)



Gianninas et al. (2015, submitted)

# Radial Velocity Follow-Up



**FLWO 1.5m**

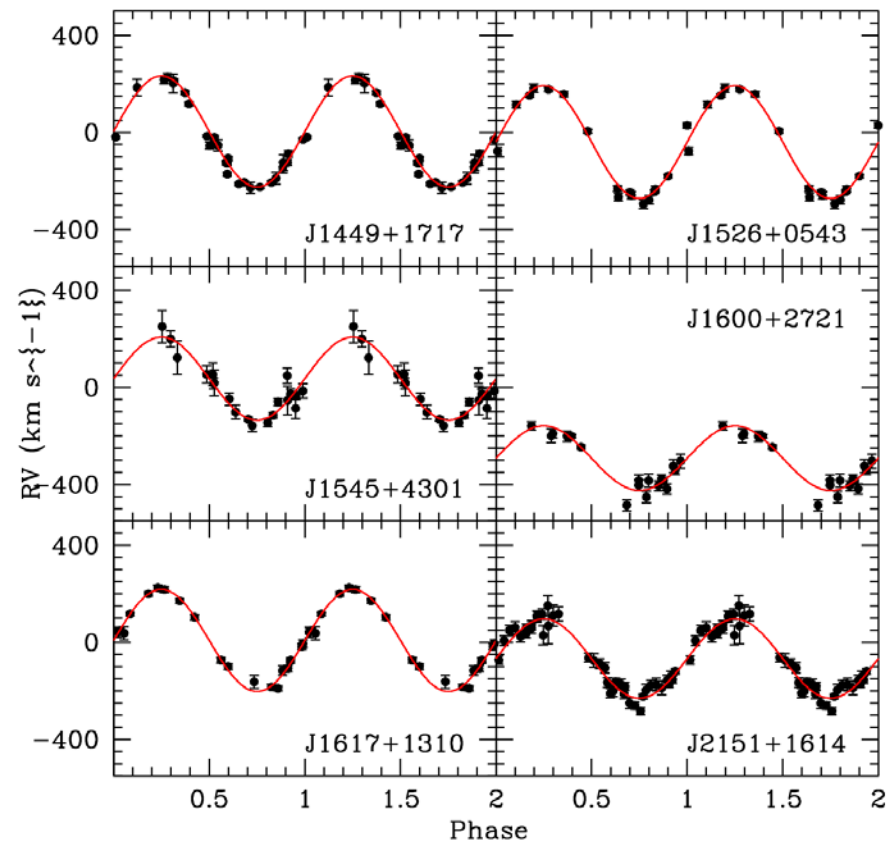
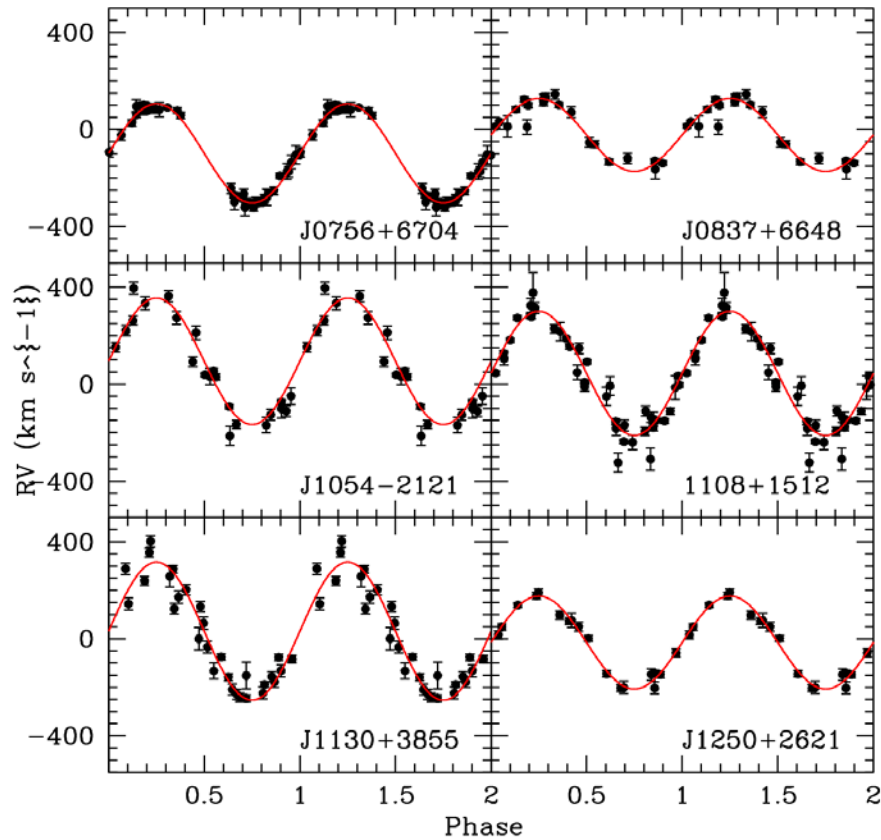


**KPNO 4m**



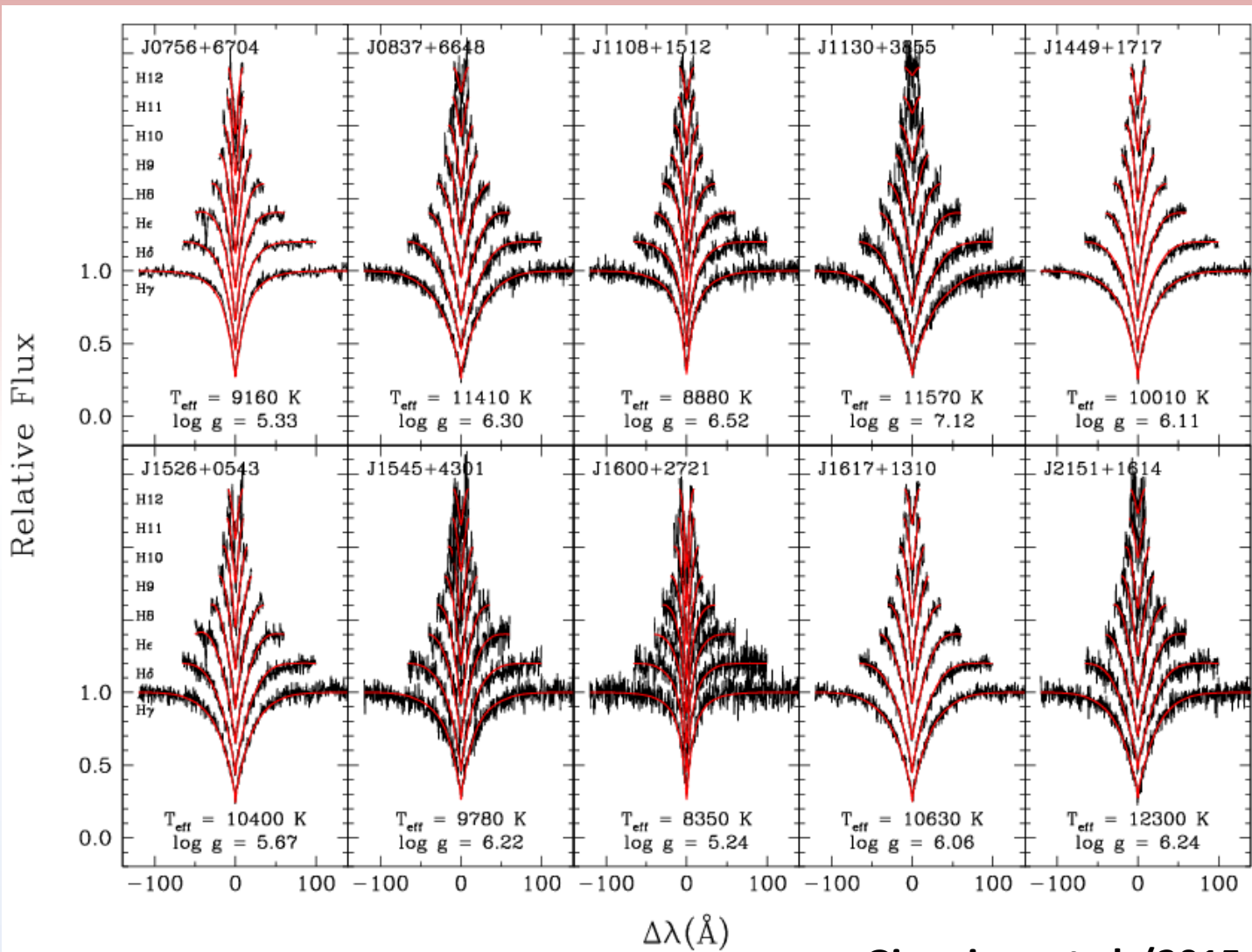
**MMT 6.5m**

# ELM VI: 12 new ELM WDs



Gianninas et al. (2015, submitted)

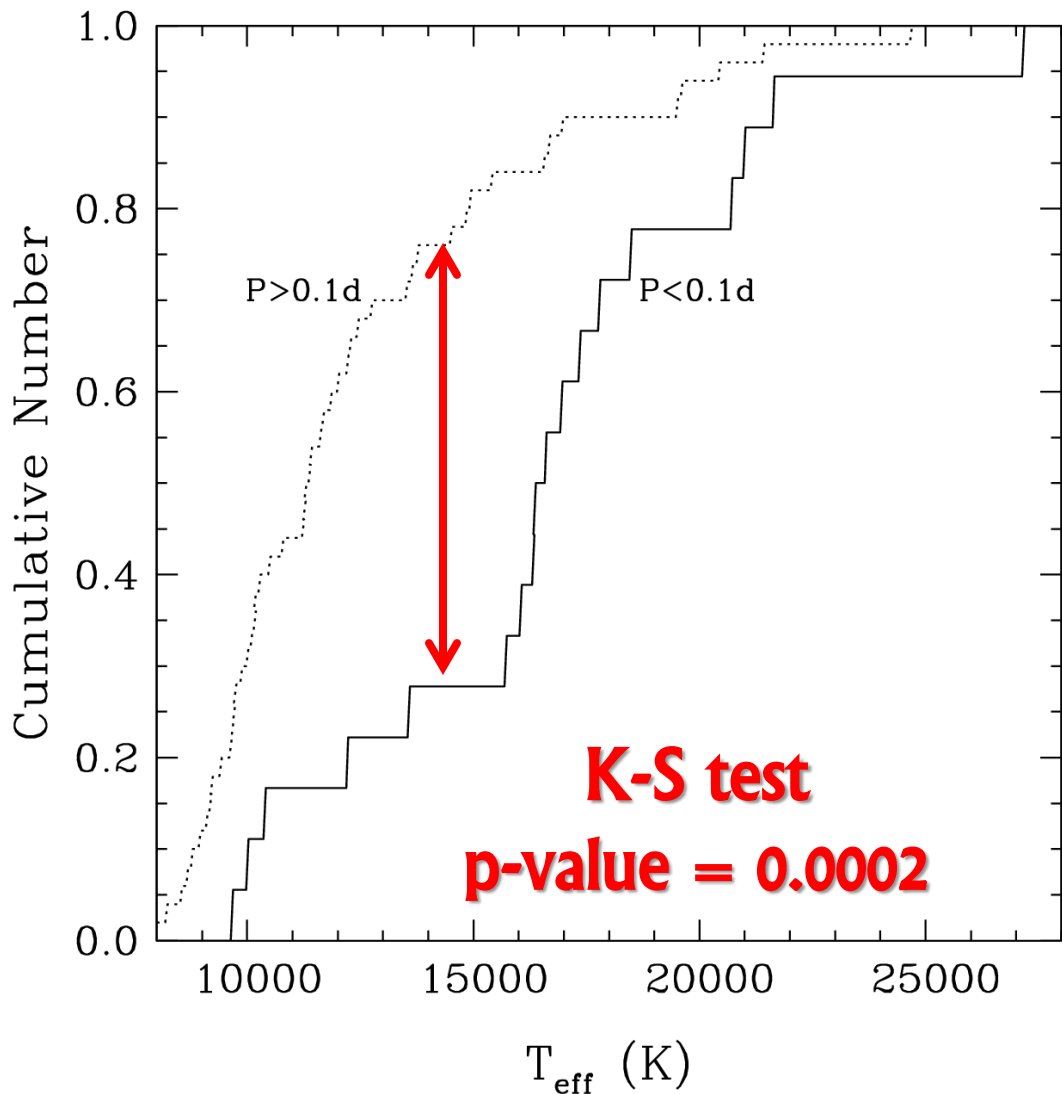
# Standard Spectroscopic Technique



Gianninas et al. (2015, submitted)

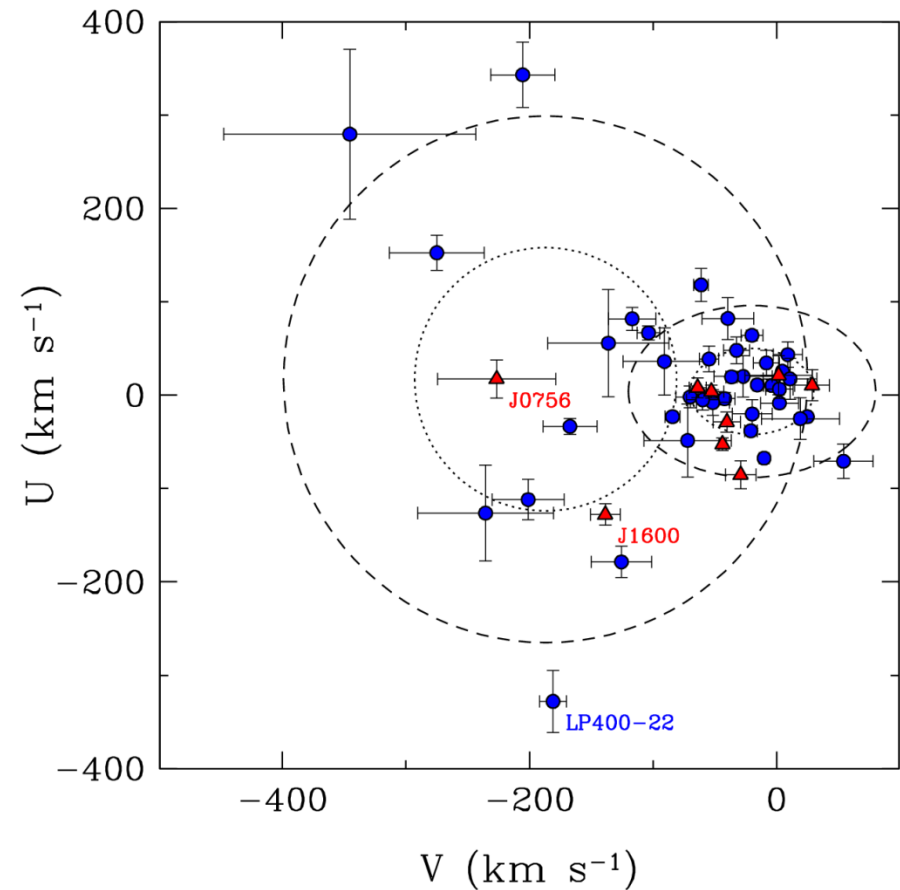
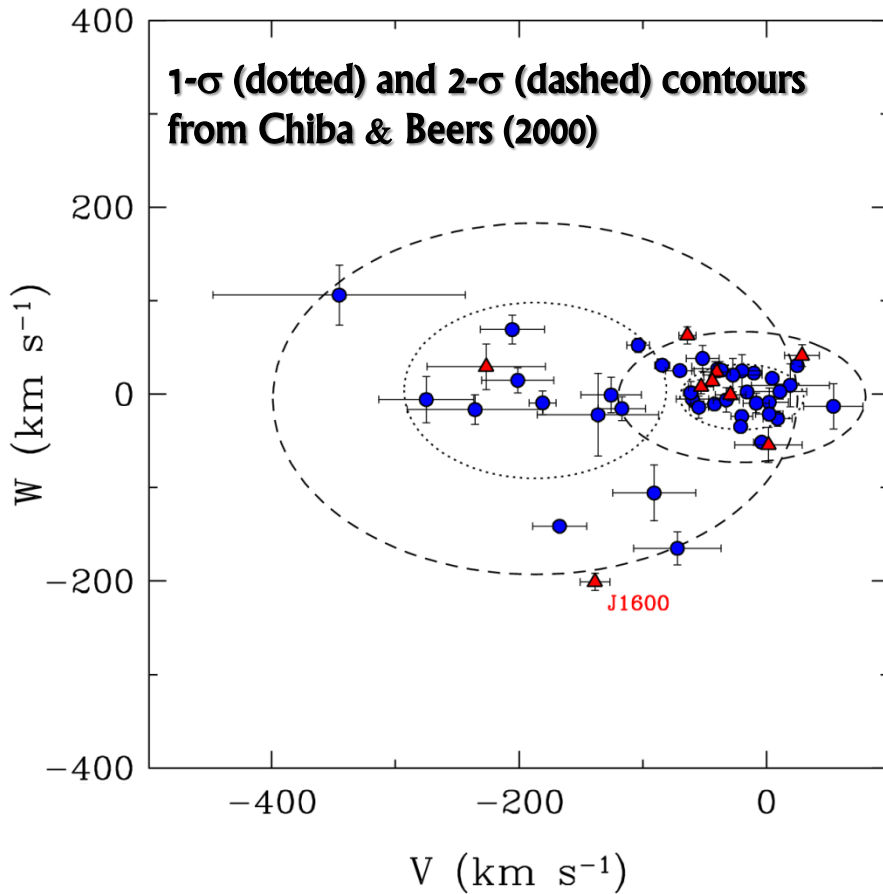
# A correlation with $T_{\text{eff}}$ is emerging

Cooler ELM WD binaries have longer periods or else they would have already merged



Gianninas et al. (2015, submitted)

# Kinematics: Space Velocities



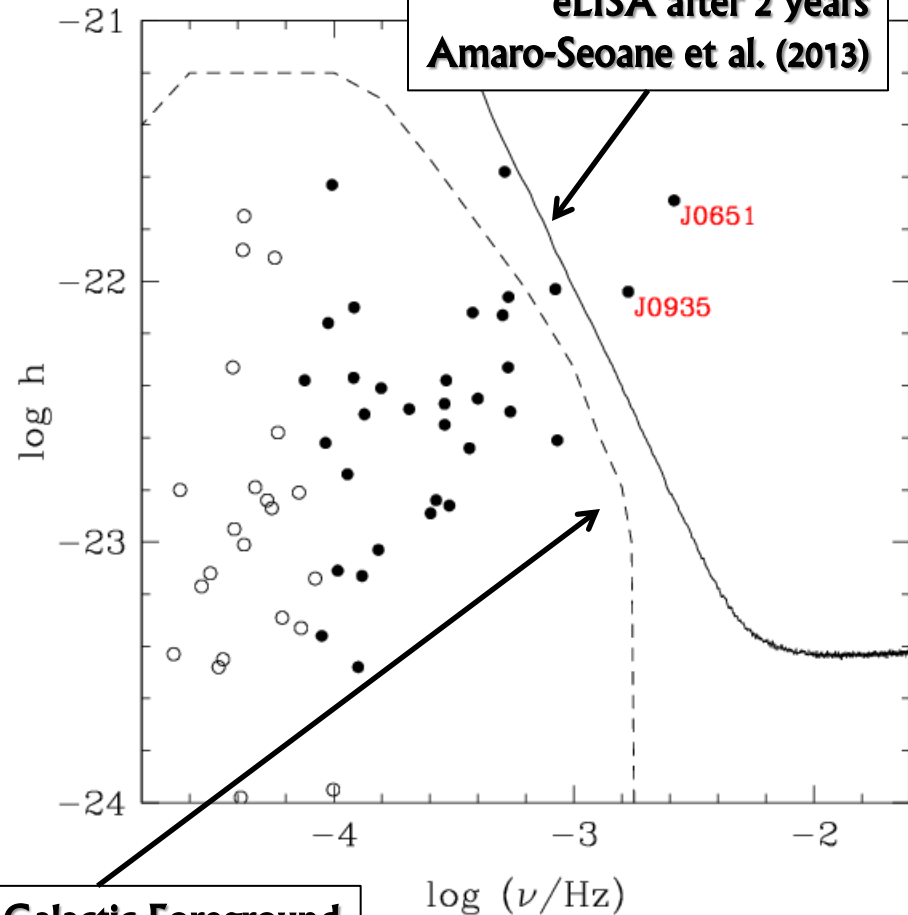
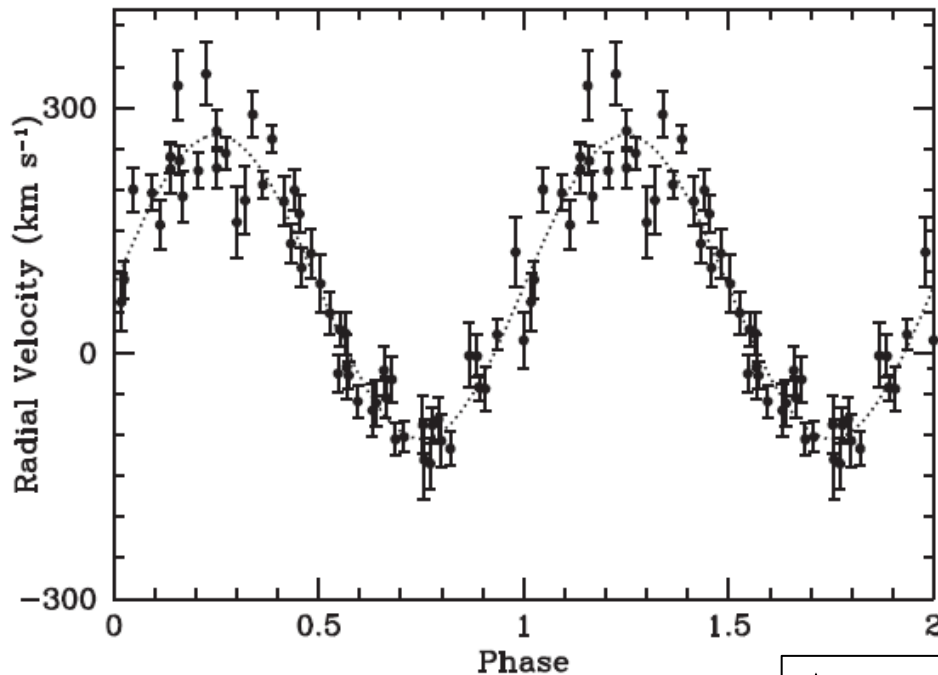
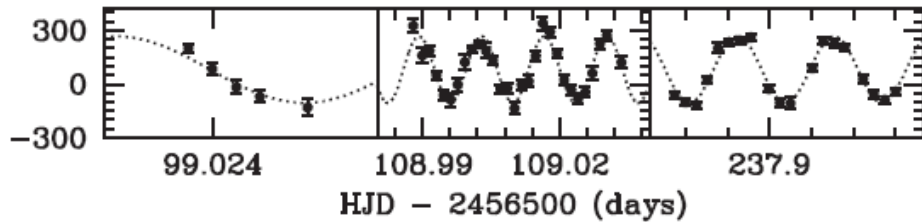
Gianninas et al. (2015, submitted)

**22 of 49 ELM WDs have kinematics consistent with halo membership**



# WD 0931+444: a new 20-min ELM WD

## Shortest period ELM after J0651!



eLISA after 2 years  
Amaro-Seoane et al. (2013)

J0651

J0935

Average Galactic Foreground  
Nelemans et al. (2001)

Kilic et al. (2015)

Kilic et al. (2014)

# Several PSR have ELM WD companions



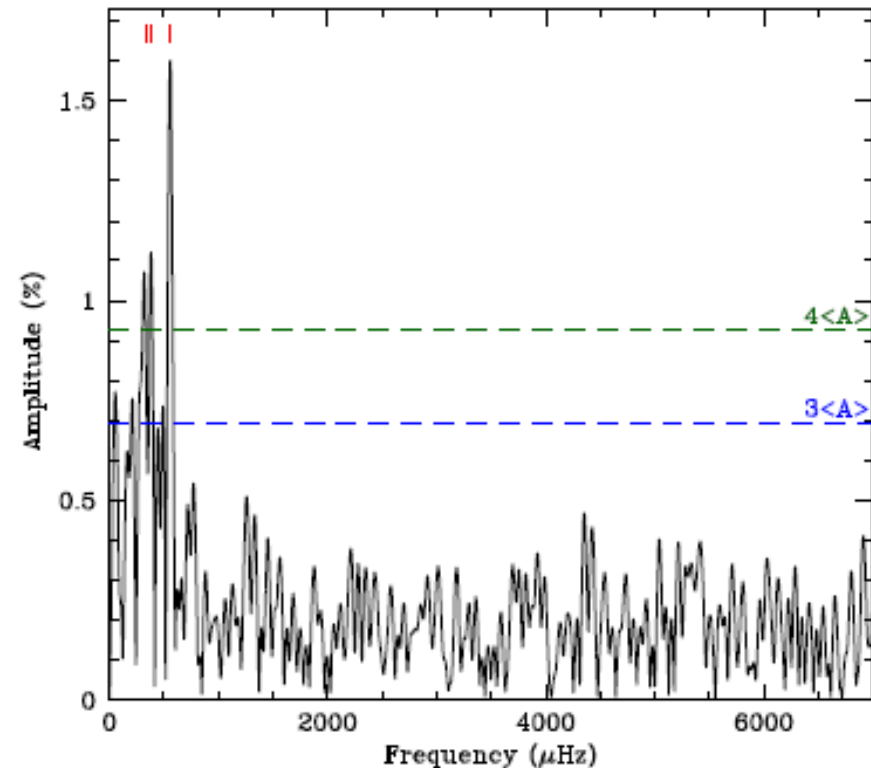
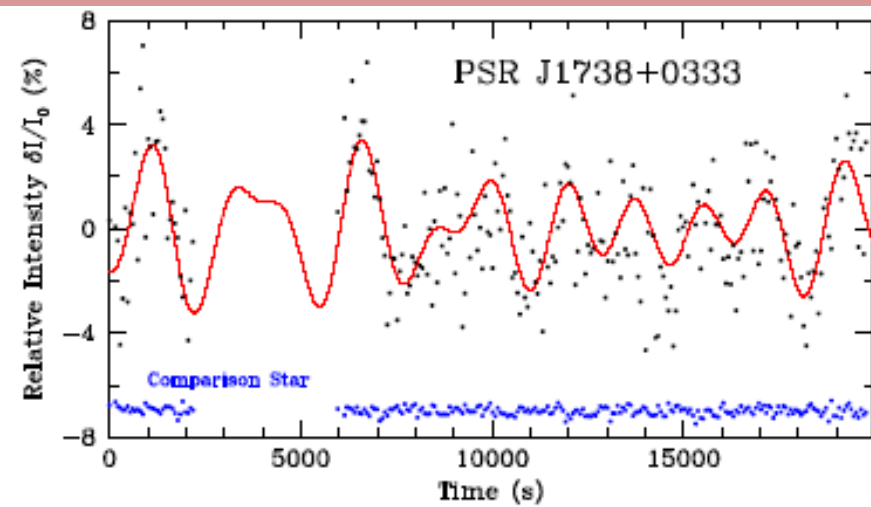
- **Several of the ELM WD companions are in the correct range of  $T_{\text{eff}}$  and  $\log g$  (i.e. instability strip) to pulsate**
- **No pulsations detected for ELM companions of PSR J1012+5307+ and PSR J1911-5958 (Steinfadt et al. 2010)**

# PSR J1738+0333

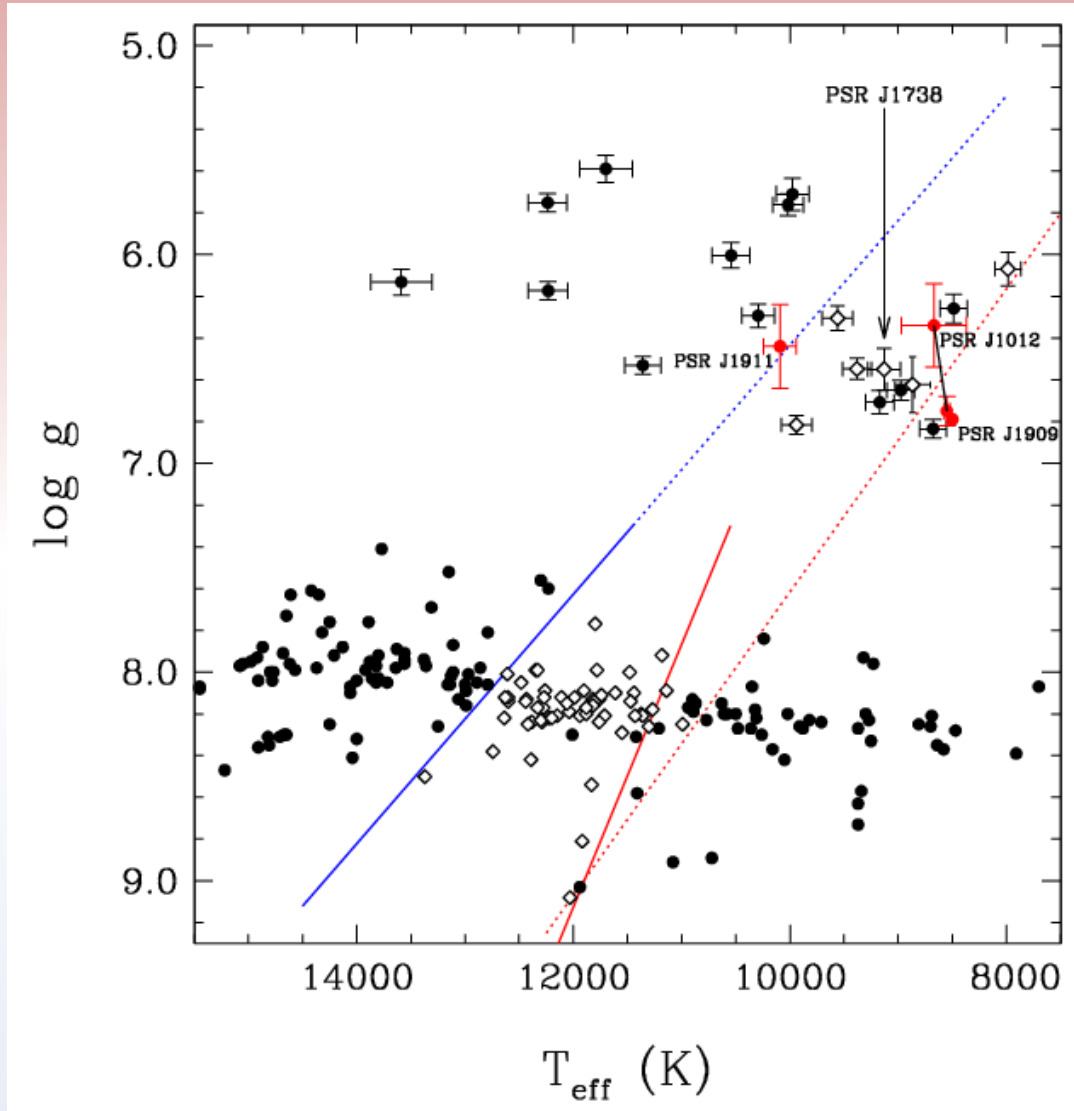
- First pulsating ELM WD companion to a PSR!

Period (s)	Frequency ( $\mu\text{Hz}$ )	Amplitude (per cent)	Phase (s)	S/N
$1788 \pm 33$	$559 \pm 10$	$1.27 \pm 0.47$	$1030 \pm 110$	6.3
$3057 \pm 99$	$327 \pm 11$	$1.22 \pm 0.47$	$10 \pm 190$	6.0
$2656 \pm 80$	$376 \pm 11$	$1.15 \pm 0.47$	$2150 \pm 170$	5.7

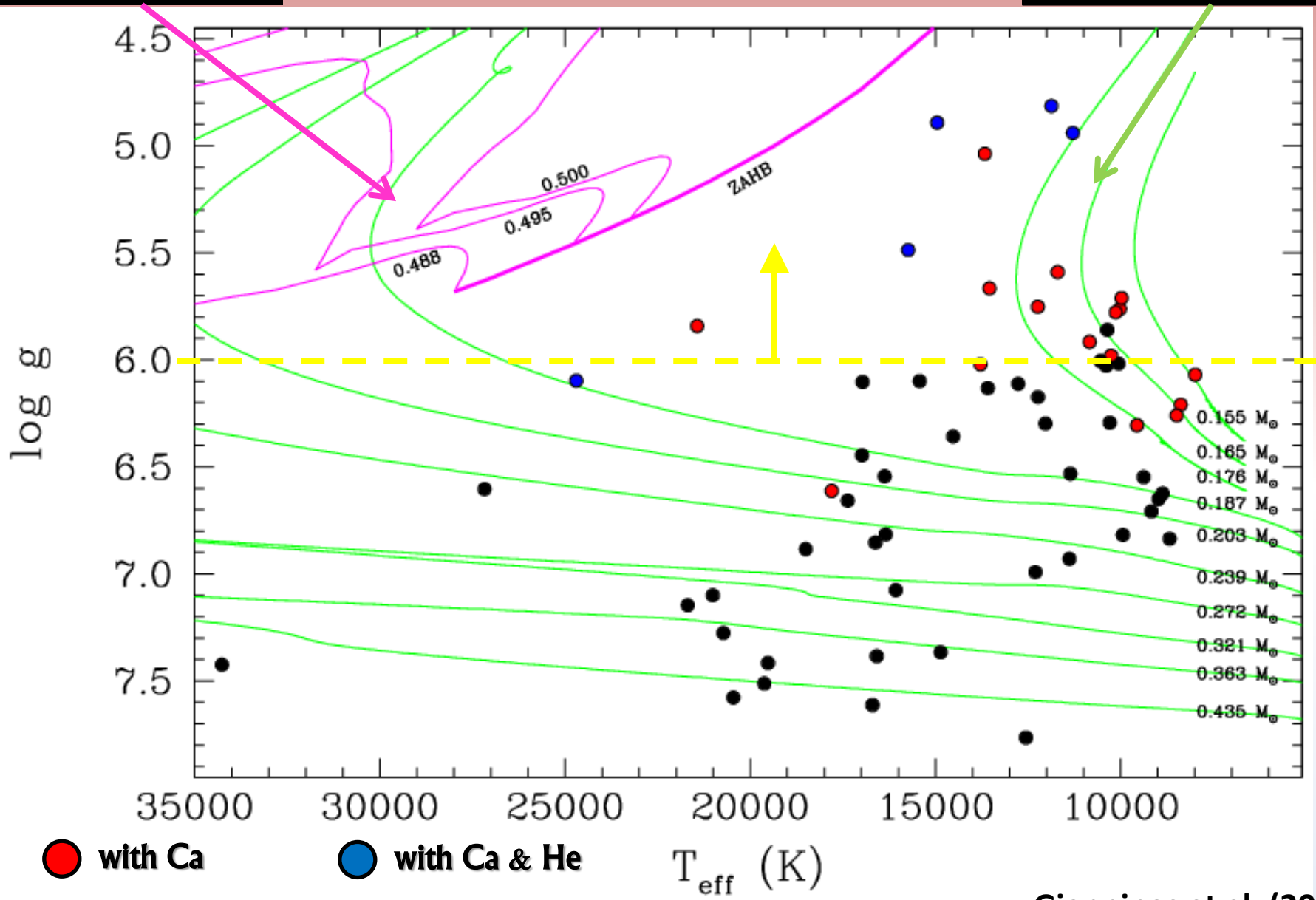
Kilic et al. (2015)



# Updated ELM WD instability strip

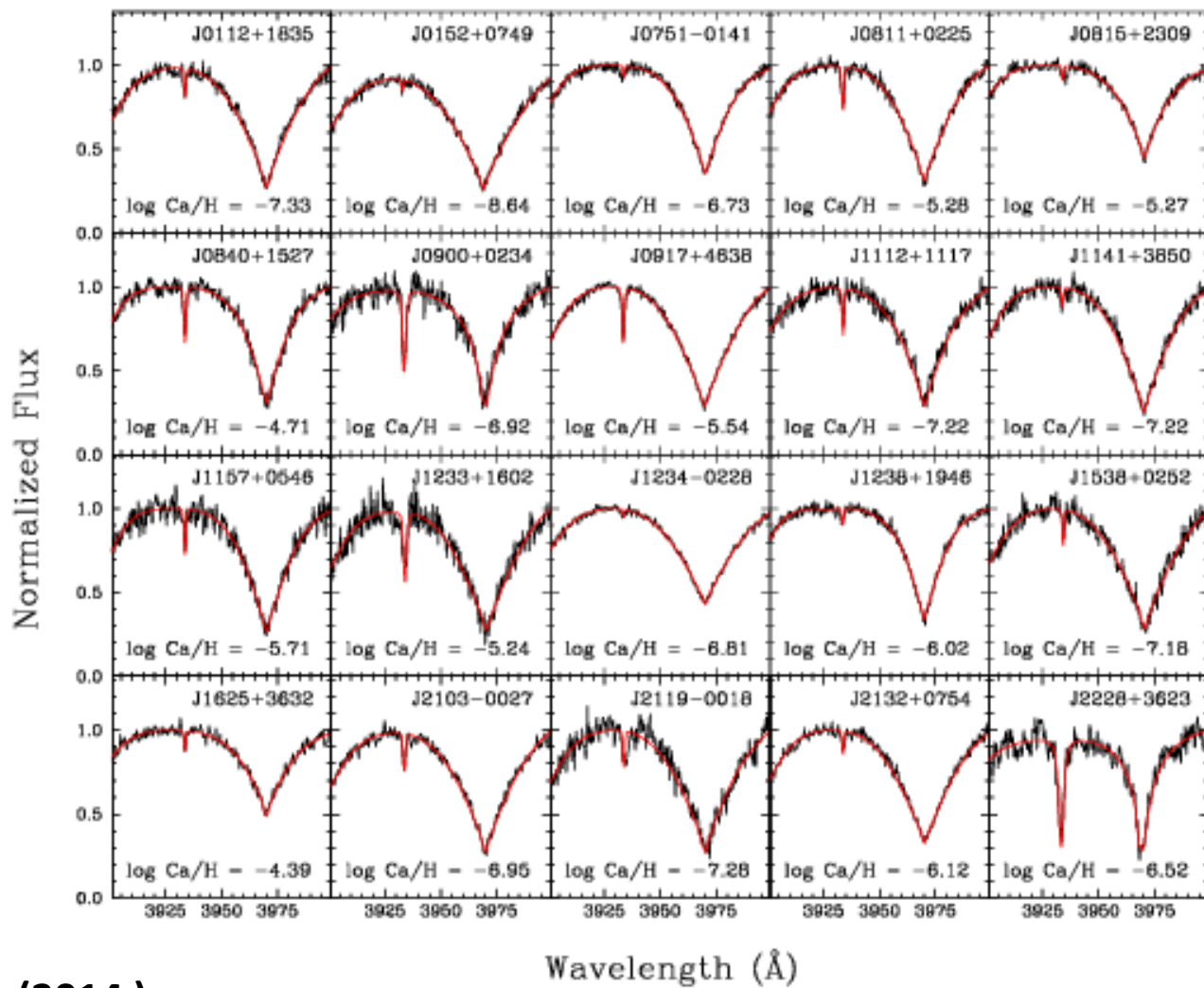


Kilic et al. (2015)



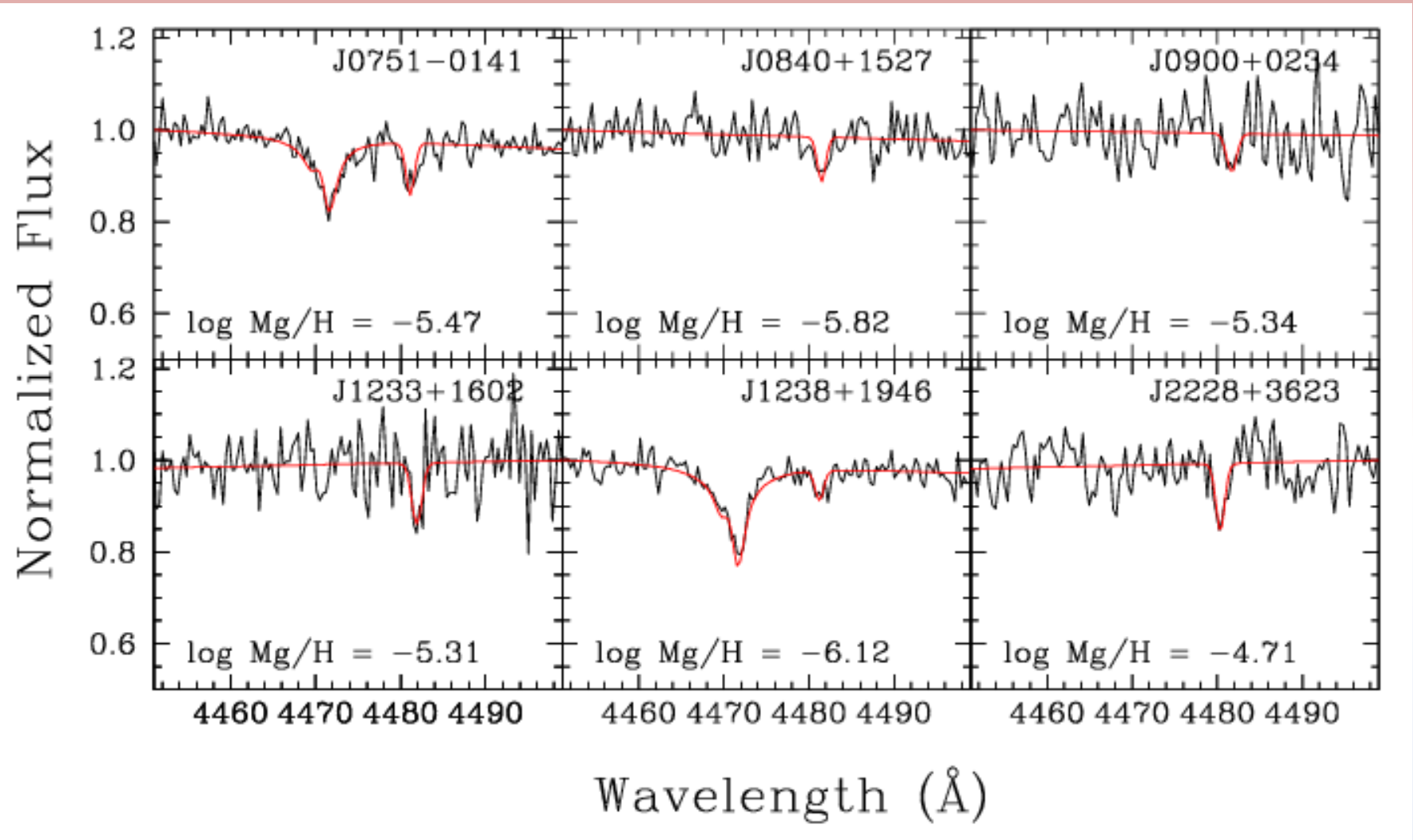
Gianninas et al. (2014)

# Many ELM WDs have Ca lines...



Gianninas et al. (2014)

... some have Mg lines...



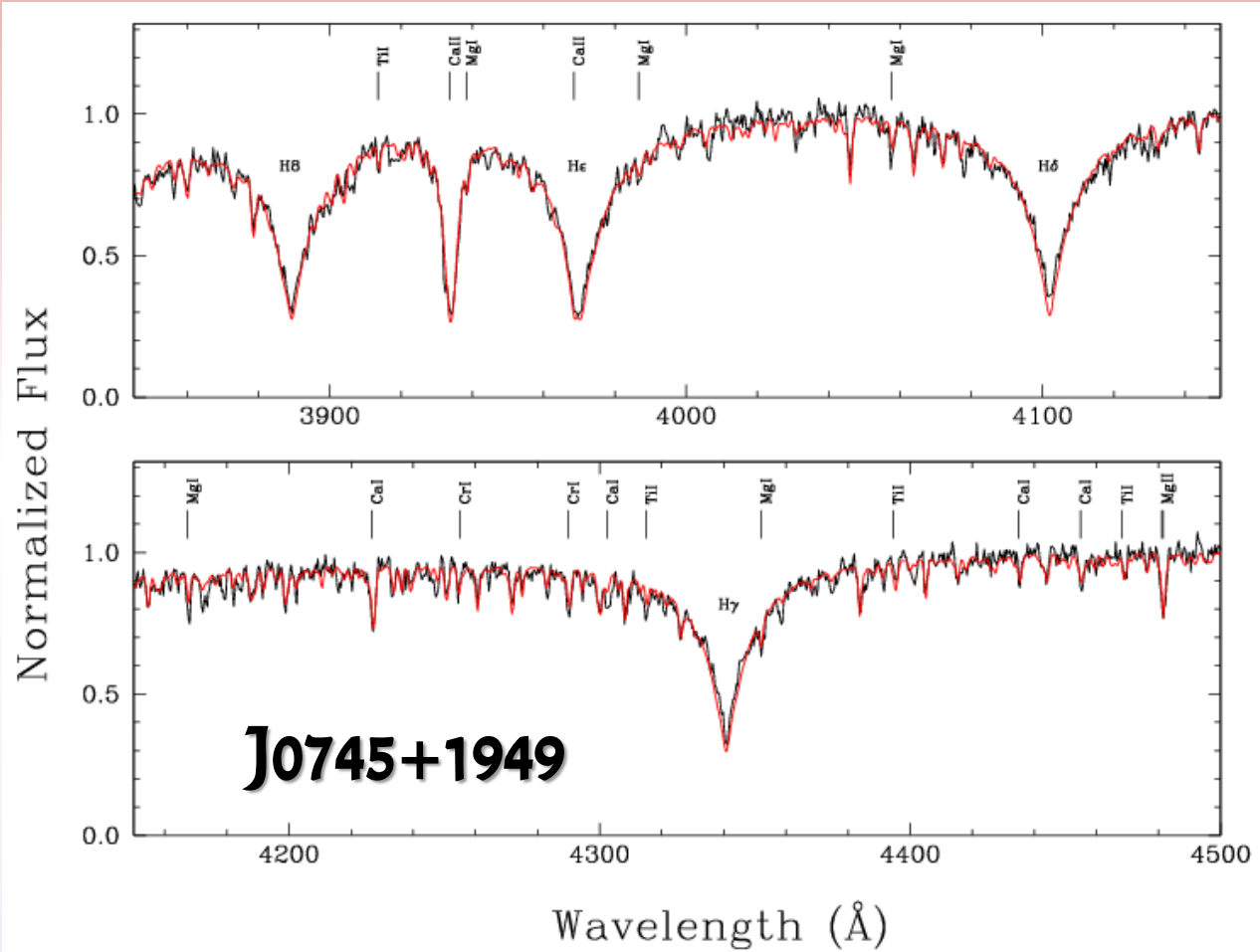
Gianninas et al. (2014)

... others have a lot more!

See also

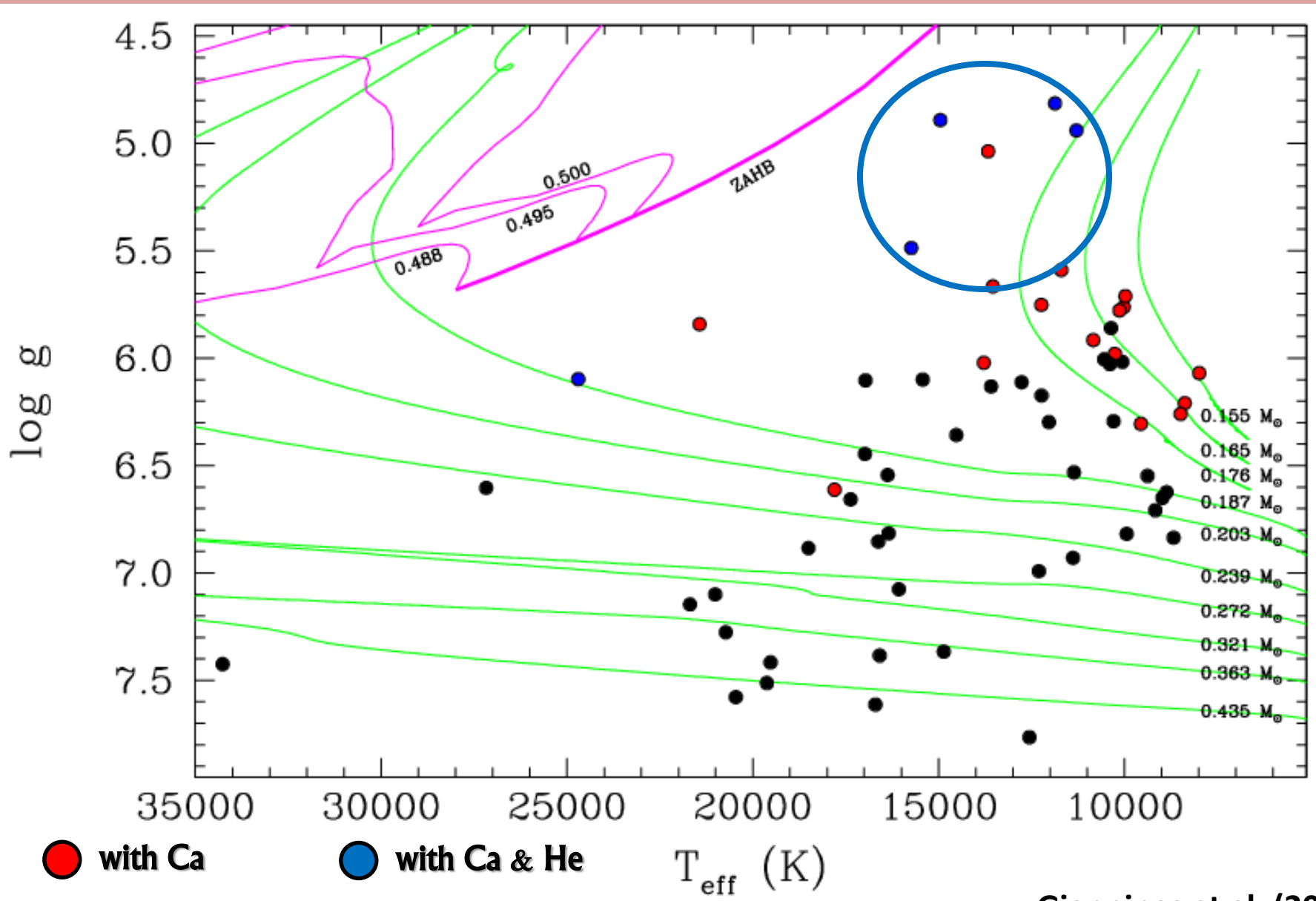
GALEX J1717+6757  
Hermes et al. (2014) +  
poster  
Vennes et al. (2011)

PSR J1816+4510  
Kaplan et al. (2013)



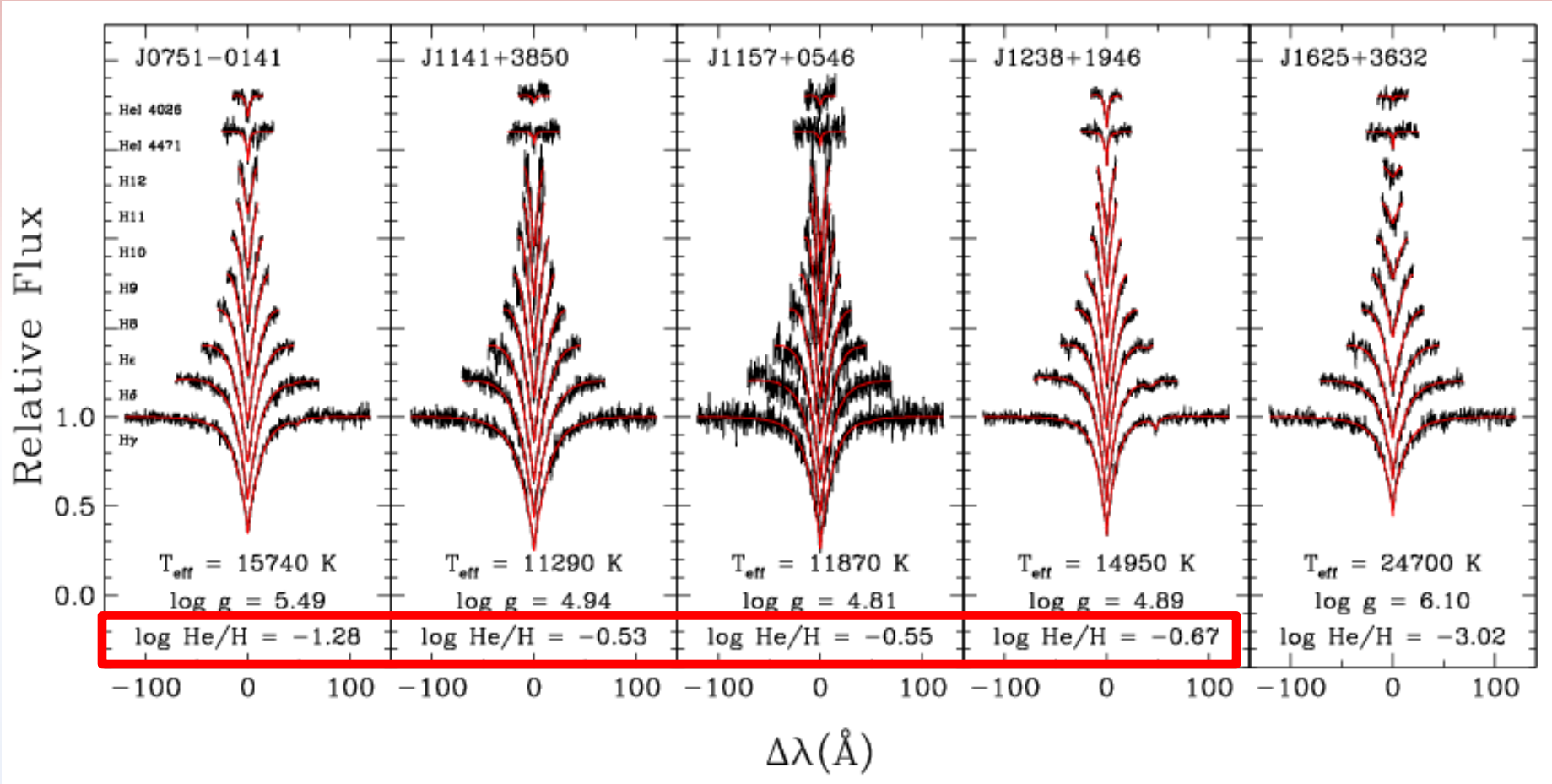
Gianninas et al. (2014)





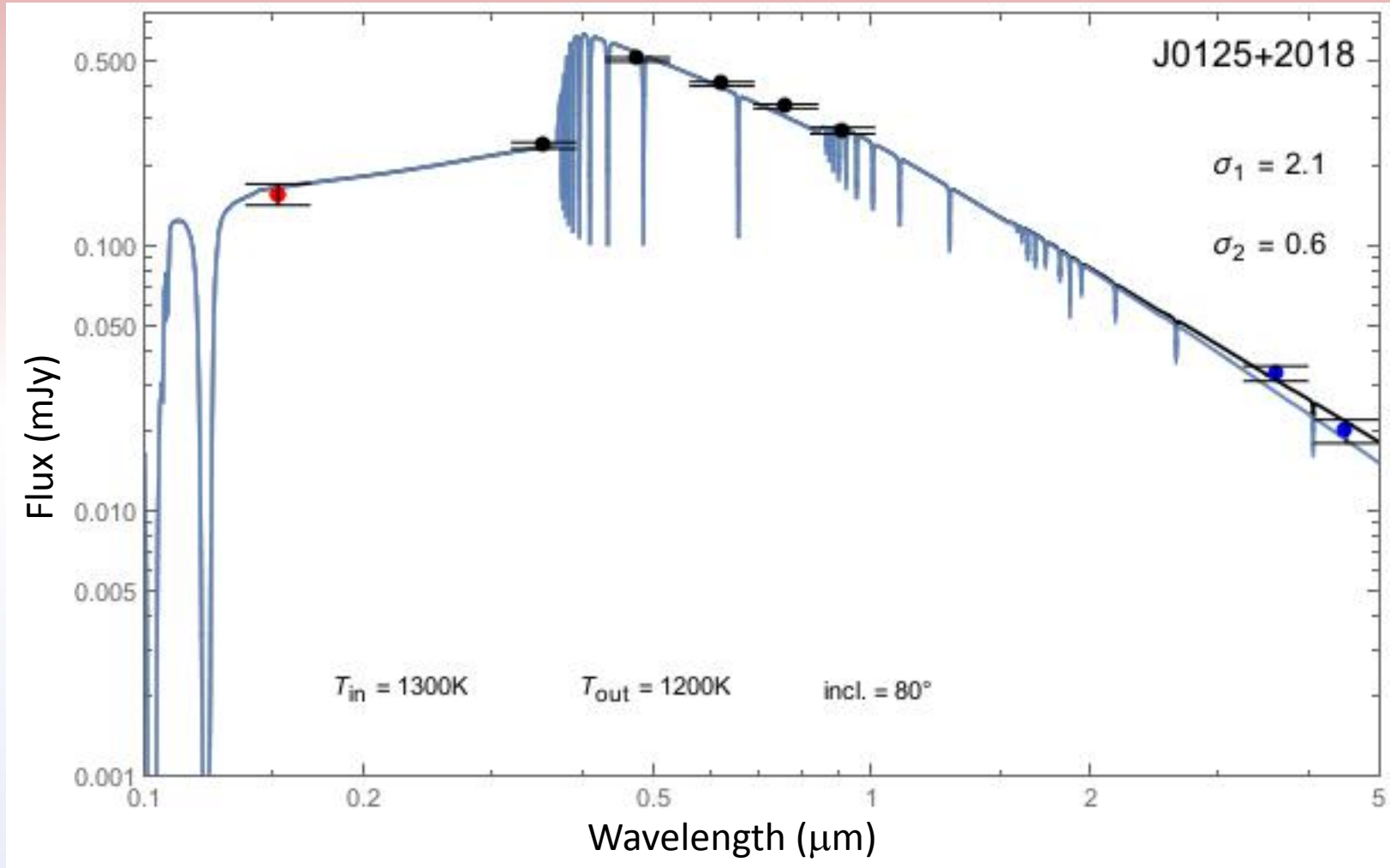
Gianninas et al. (2014)

# Some even have Helium a lot of it!

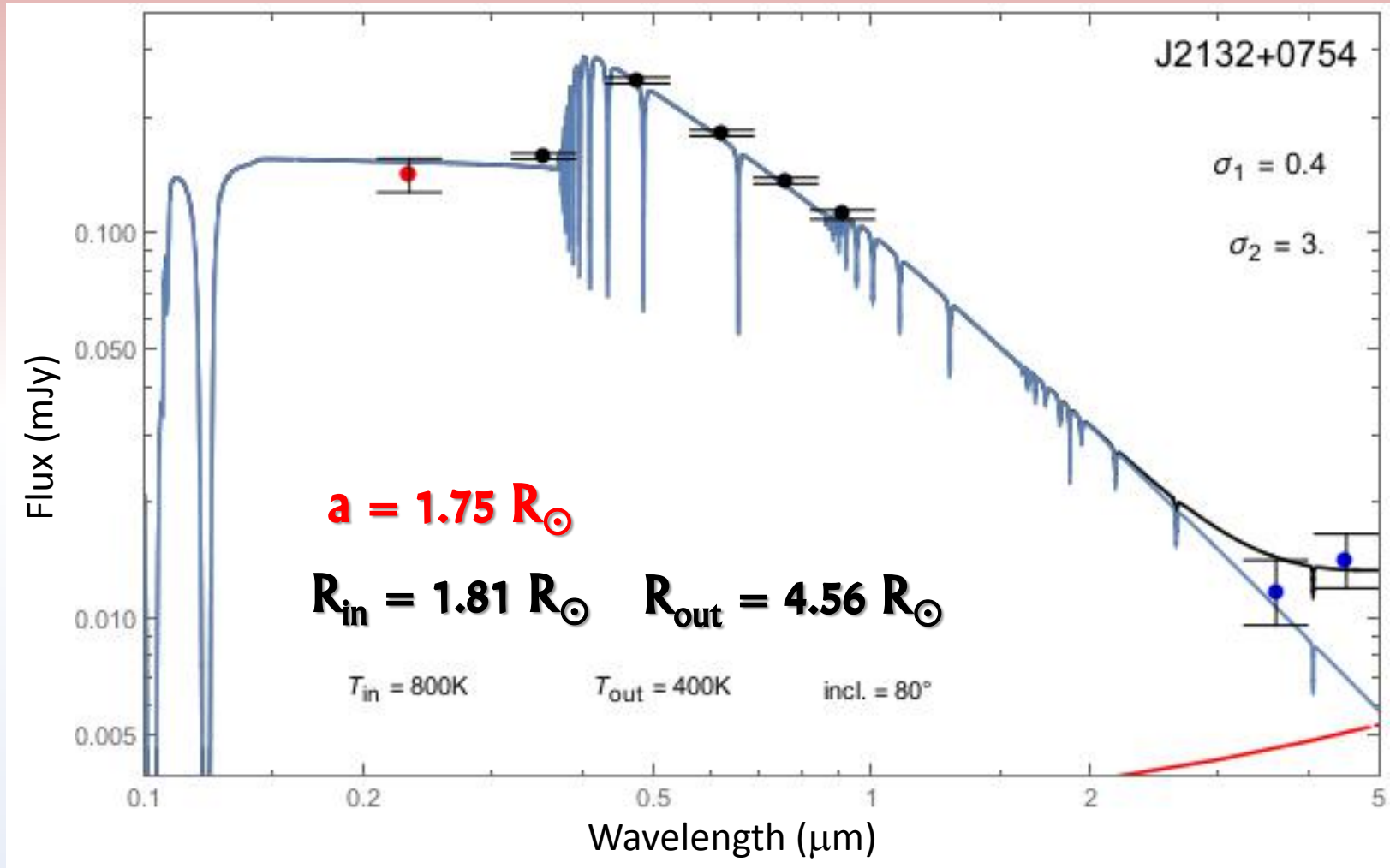


Gianninas et al. (2014)

# UV-VIS-NIR SEDs including Spitzer



# UV-VIS-NIR SEDs including Spitzer



# CONCLUSIONS

- **ELM Survey: 74 new ELM WDs and more to come**
  - Starting to be able to do some statistics + kinematics
- **Highlights:**
  - Now two gravitational wave verification sources
  - First ever PSR+ pulsating ELM WD
- **Metals:**
  - Ca (Mg and more) in all ELMs with  $\log g < 6.0$
  - Presence of He could be sign of recent shell flash
  - No excesses in Spitzer data (except one...?)
- **Wish list:**
  - Constrain parameters of the secondary components