

A new HW Vir binary from the Palomar Transient Factory

PTF1 J072455.75+125300.3

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1 HW Vir systems

2 PTF 072456+125301

- Observations
- Analysis
 - Spectroscopic analysis
 - Lightcurve analysis

3 Summary & Outlook

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HW Vir Objects

- sdB + dwarf companion
- eclipsing binary system
- 17 known systems
- characteristic lightcurve (reflection effect)
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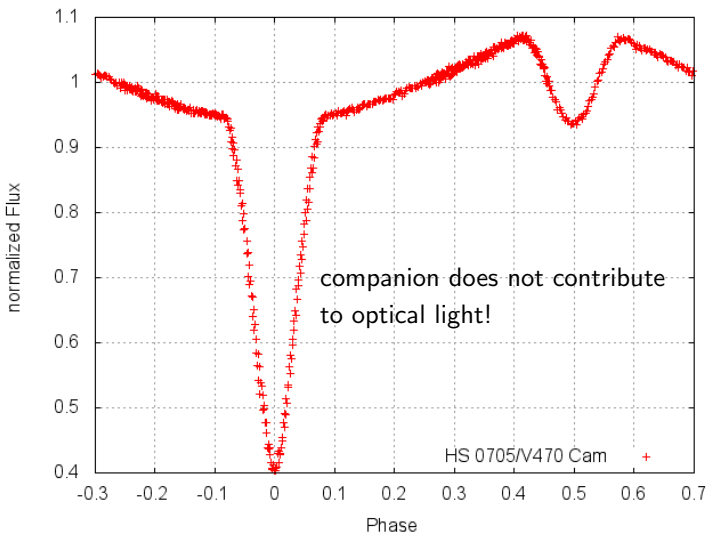


Figure : HS 0705, Data provided by H.Drechsel

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Detected at the **Palomar Transient Factory (PTF)**

- Palomar Observatory (Caltech)
- Wide field survey
- Search for optical transients and variable sources
- 48" telescope: robotic survey-telescope
- 60" telescope: follow-up observations

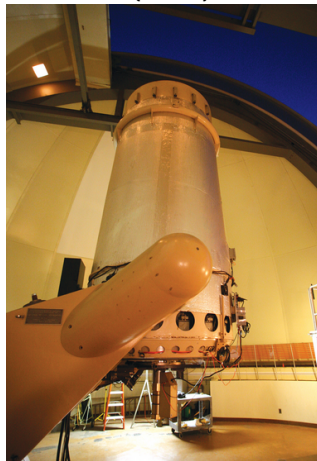


Figure : <http://www.astro.caltech.edu/palomar/about/telescopes.html>

Obtaining the data...

● Photometry

- seven lightcurves, each ~ 150 datapoints
- blue filter (400 nm to 550 nm)
- mostly covered more than one orbit $\rightarrow P=0.09980(25)$ d
- six lightcurves used for analysis
- phased, normalized and merged to one dataset

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● Spectroscopy

- 38 spectra
- Palomar 200" and Double Spectrograph (DBSP)
- $R \sim 1500$

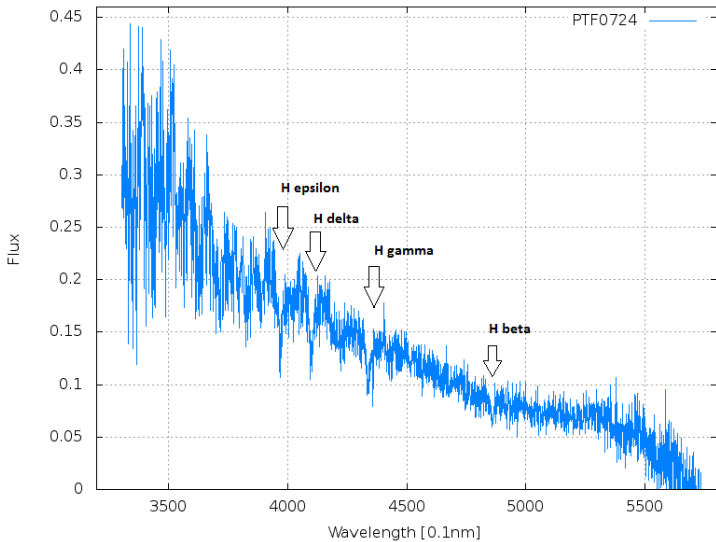
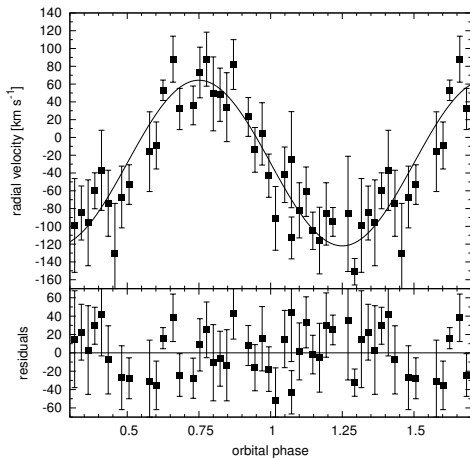


Figure : Spectrum of PTF0724 with hydrogen lines

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Radial Velocity Curve



- lines: H β , H γ , H δ , (He)
- circular orbit
- $P=0.09(85)$ d
- semi-amplitude: $95.7 \pm 8.1 \text{ km s}^{-1}$
- system velocity: $-26.7 \pm 5.5 \text{ km s}^{-1}$
- mass function: $\rightarrow q = 0.33$

Figure : Radial velocity curve and residuals

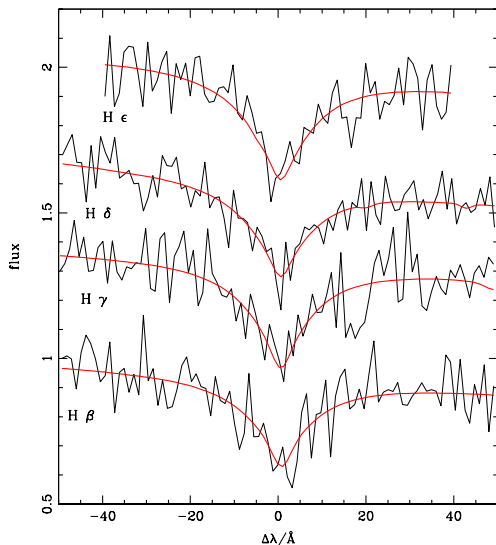


Figure : Example fit to several Balmer lines of an individual spectrum to determine the radial velocities.

Atmospheric Parameters

Determination of atmospheric parameters by fitting synthetic spectra to Balmer- and He-lines of coadded spectra.

Synthetic spectra: LTE (Local Thermodynamical Equilibrium) model atmospheres with solar metallicity and line blanketing

- $T_{\text{eff}} = 33900 \pm 350 \text{ K}$
- $\log g = 5.74 \pm 0.08$
- $\log y = -2.02 \pm 0.07$

Comparison with other sdB systems...

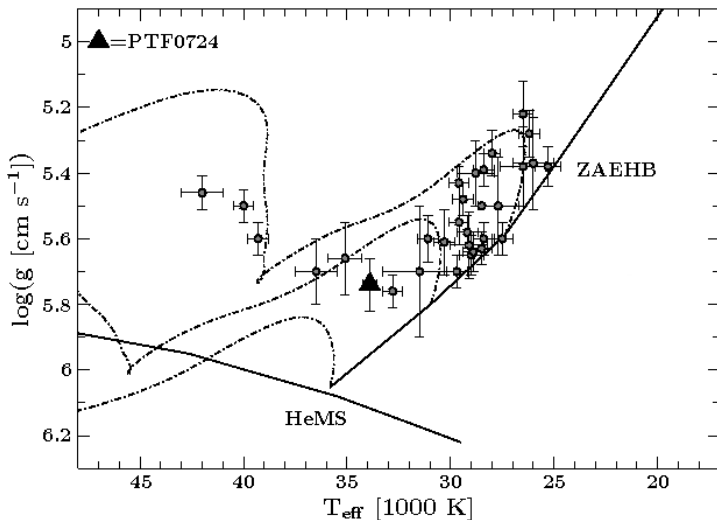


Figure : Position of PTF0724 in a (T_{eff} , $\log g$) diagram in comparison to other known sdB stars in close binary systems.

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Basic information

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 - vary mass ratio around $q=0.33$
- ~ 2000 sets of start parameters
- sum of squared residuals of all datapoints → estimation of quality

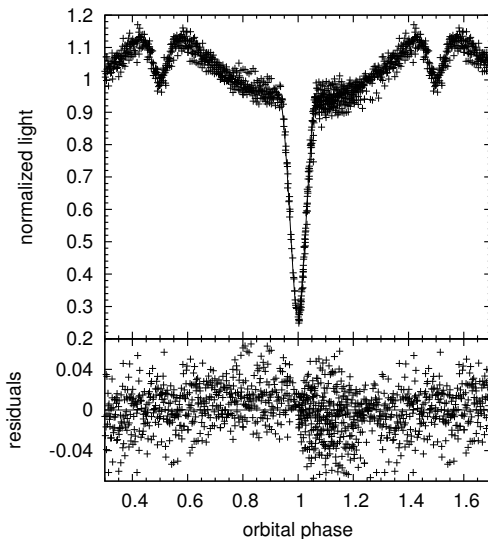


Figure : Lightcurve and best-fit. Bottom panel: residuals

Fixed parameters:		Fitted parameters:	
$T_{\text{eff}}(1)$	33900 K	i	$83.56 \pm 0.3^\circ$
g_1	1.0	$T_{\text{eff}}(2)$	$3300 \pm 300 \text{ K}$
g_2	0.32	A_2	1.8 ± 0.2
x_1	0.190	Ω_1	5.544 ± 0.20
A_1	1.0	Ω_2	2.885 ± 0.15
δ_2	0.0	x_2	0.510 ± 0.05
l_3	0.00	$\left(\frac{L_1}{L_1+L_2}\right)$	(0.99978 ± 0.00064)
		(δ_1)	(0.0006 ± 0.0003)
		r_1	0.192 a
		r_2	0.218 a

Table : Parameters from the lightcurve analysis

PTF1J072456+125301		
i	°	83.56 ± 0.30
M_{comp}	$[M_{\odot}]$	0.155 ± 0.020
a	$[R_{\odot}]$	0.766 ± 0.041
R_{sdB}	$[R_{\odot}]$	0.1488 ± 0.007
R_{comp}	$[R_{\odot}]$	0.165 ± 0.008
$\log g(\text{sdB, phot})$		5.760 ± 0.015
$\log g(\text{sdB, spec})$		5.75 ± 0.08

Table : Adopted stellar parameters for the components of PTF0724.

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Future tasks:

- better spectra
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- Pulsations (similar to 346 Hya stars) ?

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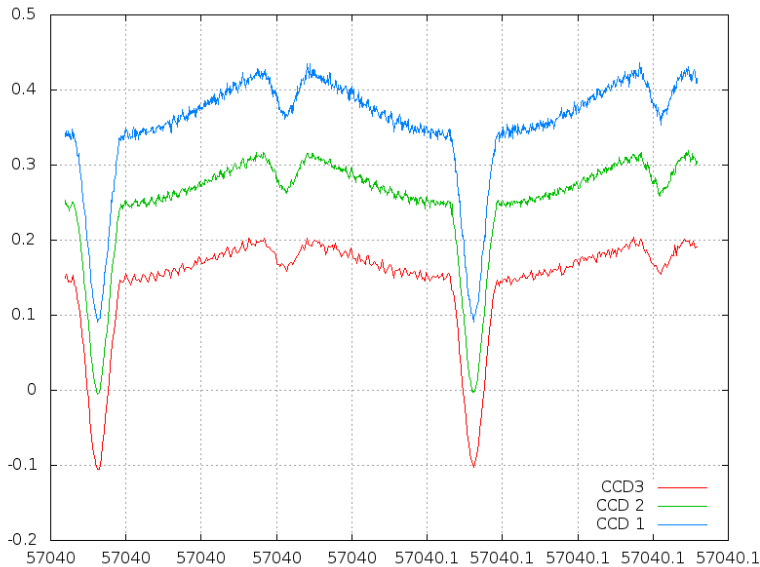


Figure : Ultracam light curve of PTF0724

Thank you for your attention!
Any questions?