

Testing eccentricity pumping processes in wide sdB binaries with MESA

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Seventh meeting on hot subdwarfs and related objects



sdB observing campaign

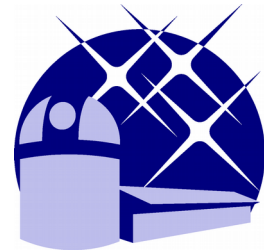


HERMES @ Mercator



- 8 targets
- 6 years of monitoring
- ~50 spectra/target

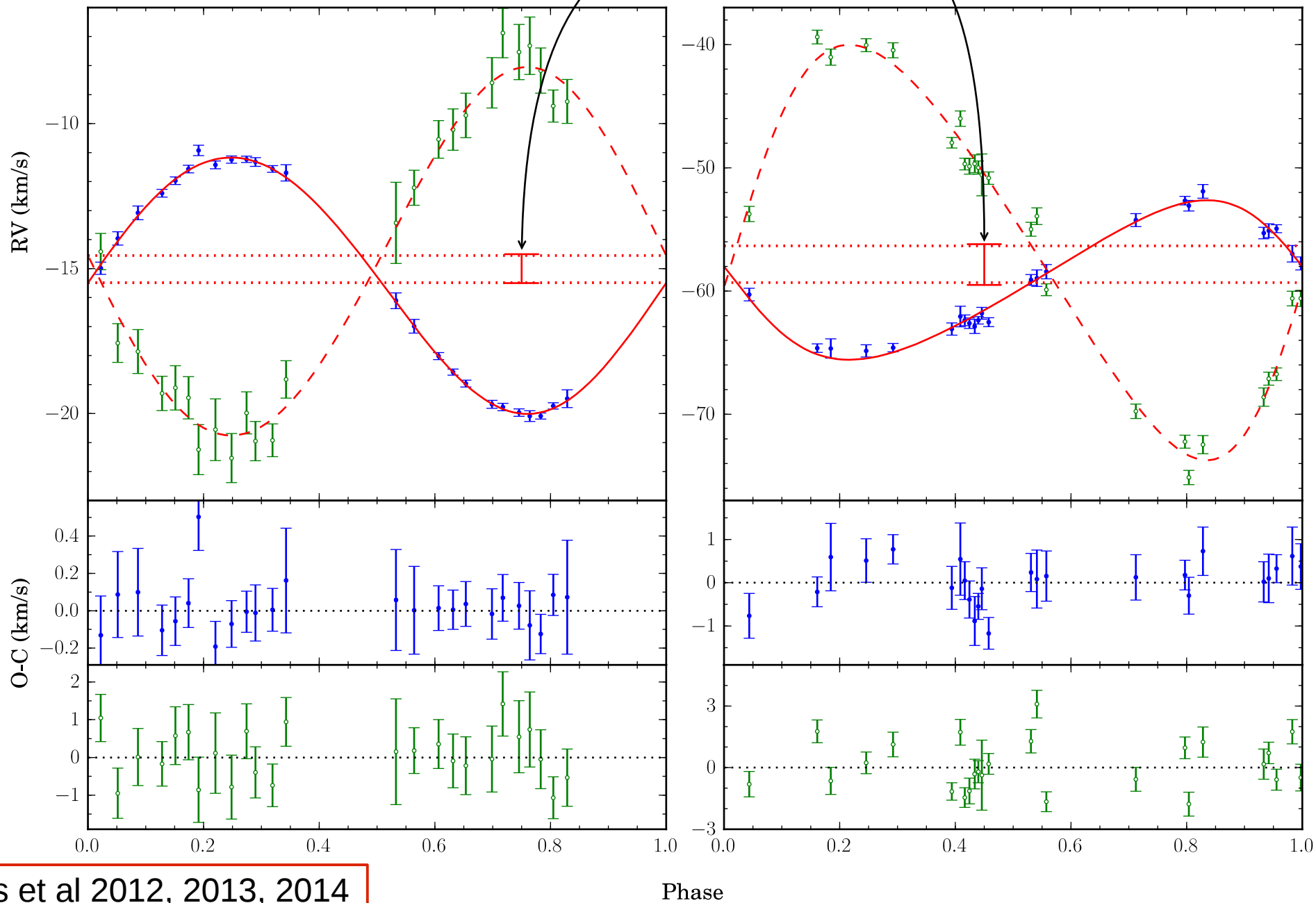
Radial velocity curves



Gravitational Redshift

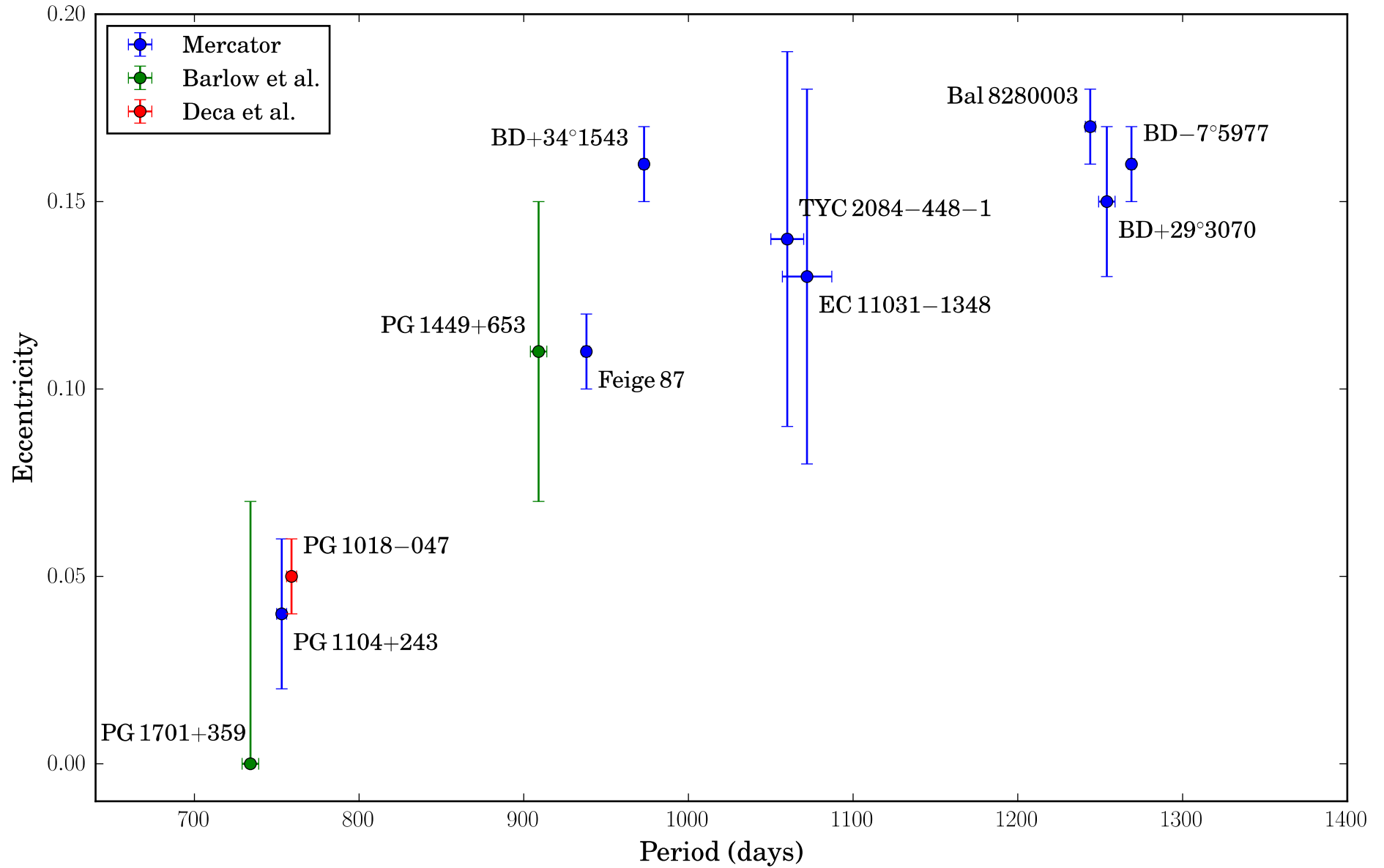
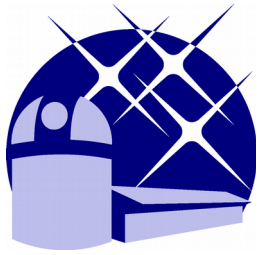
PG 1104+243

BD +29°3070

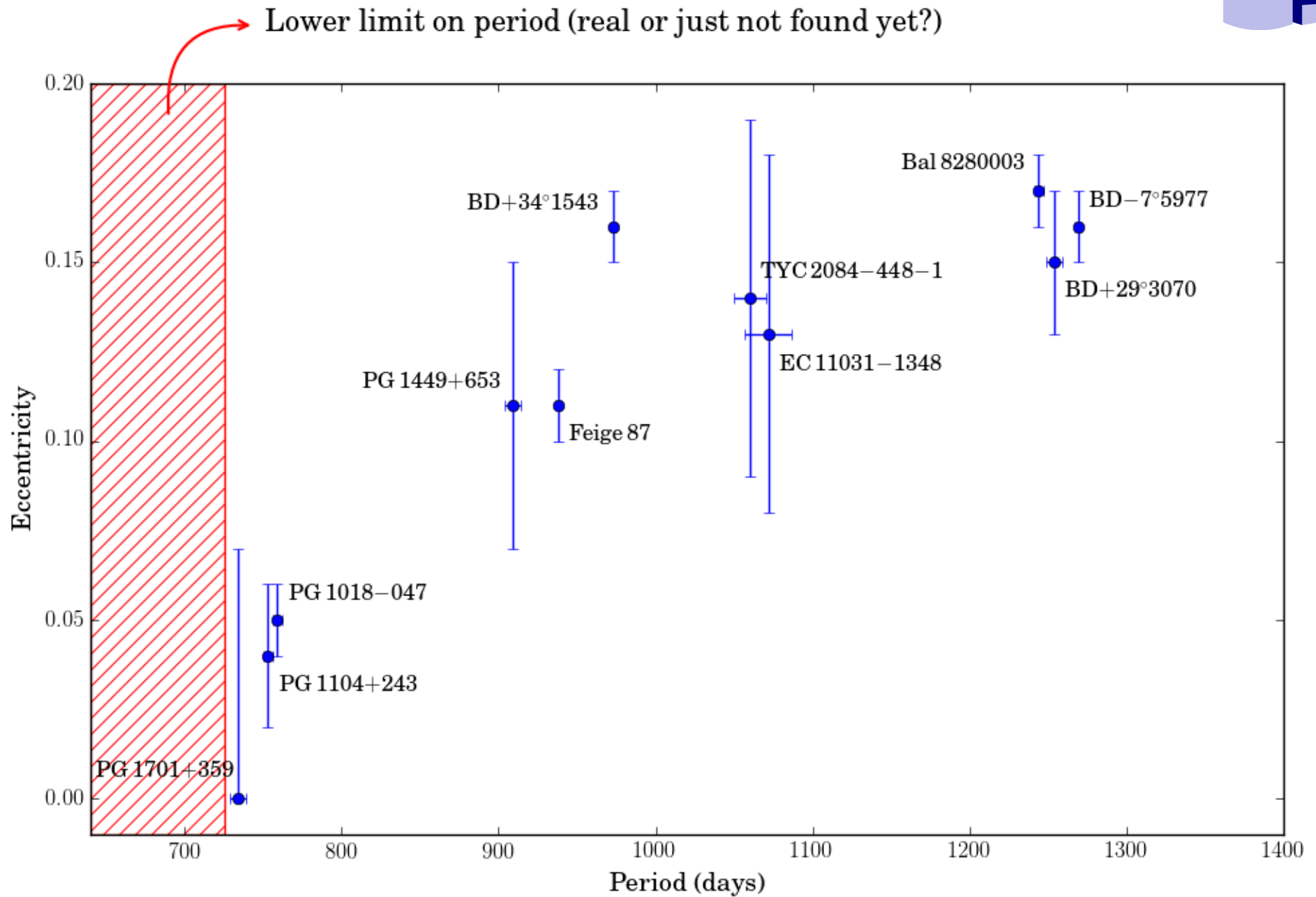
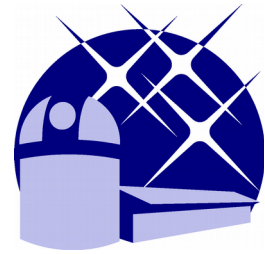


Vos et al 2012, 2013, 2014

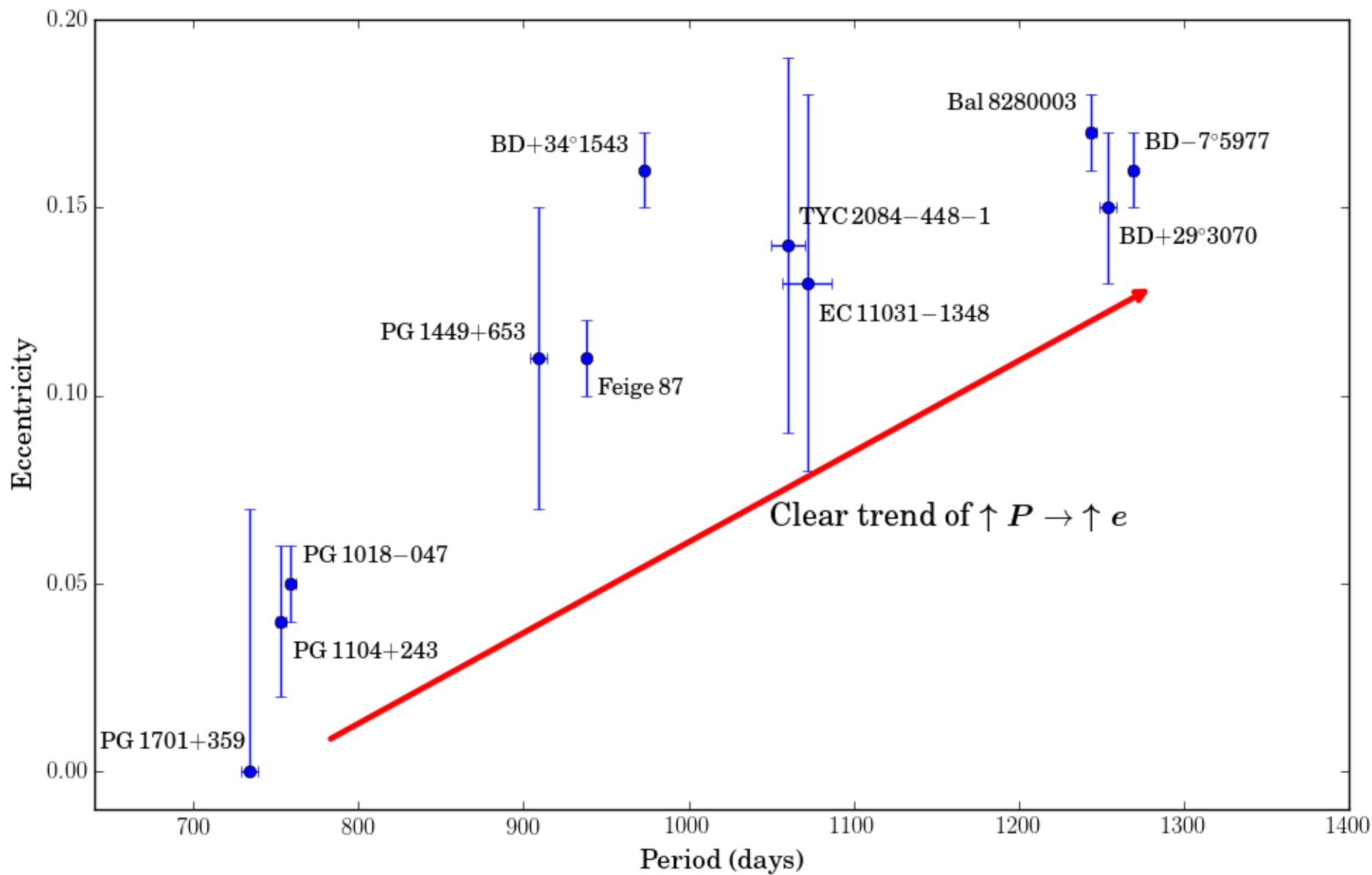
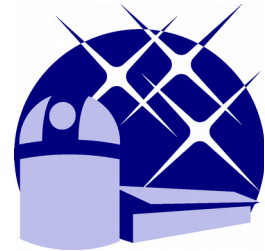
Observed period - eccentricity



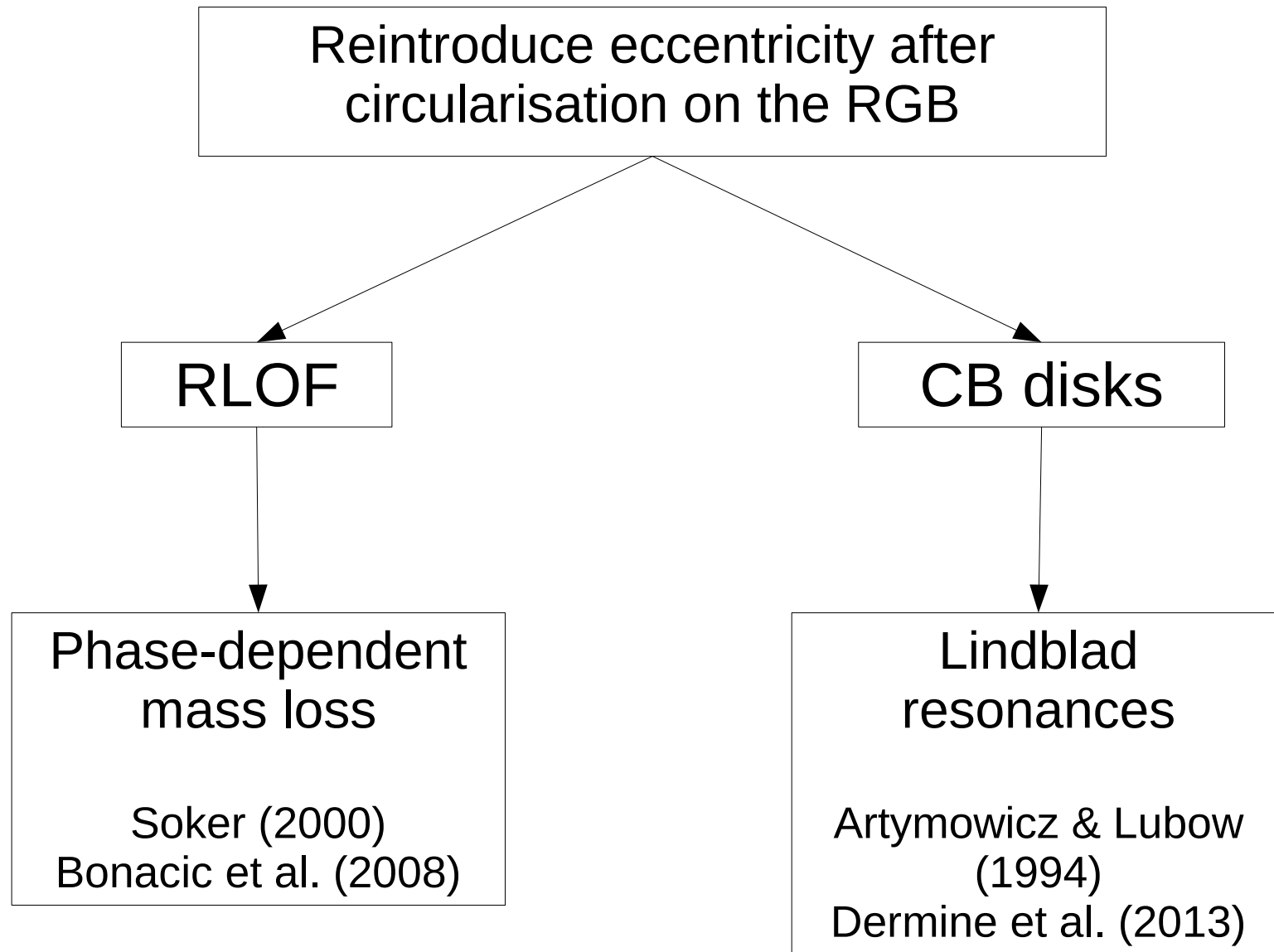
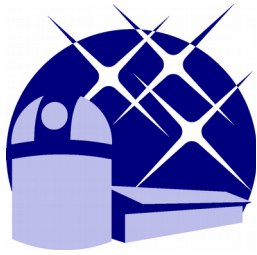
Observed period - eccentricity



Unexpected eccentricity

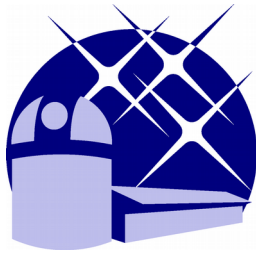


Eccentricity Pumping



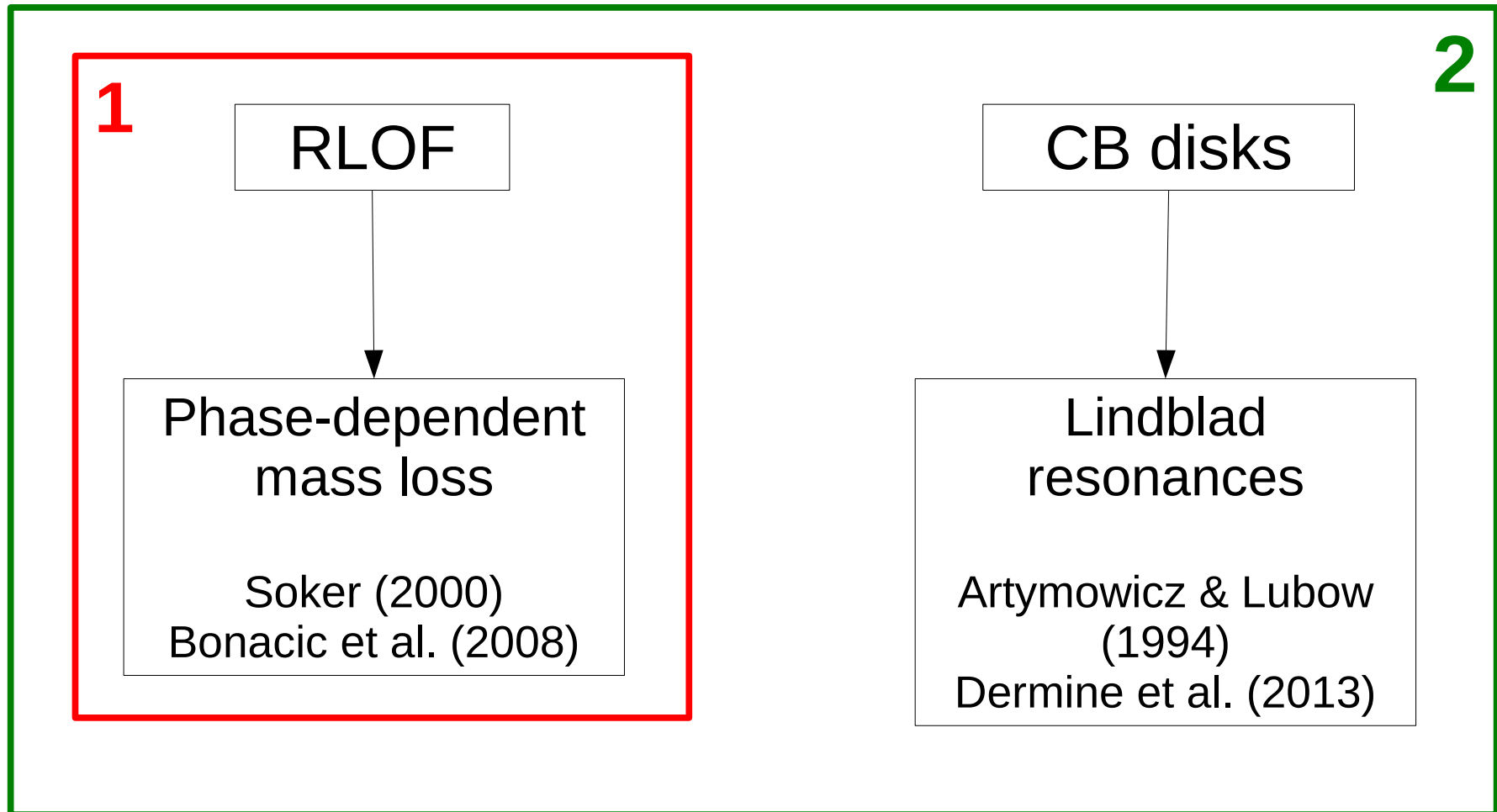
Needs minimum eccentricity $e_{\min} = 0.001$

Model setup

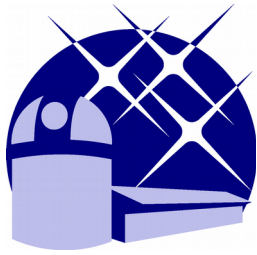


1) Only models with phase dependent RLOF

2) Models with phase dependent RLOF and a CB disk

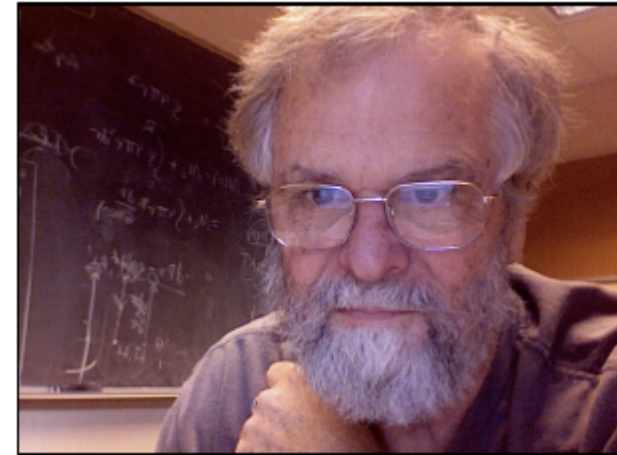


MESA



Modules for Experiments in Stellar Astrophysics

MESA is a state-of-the-art,
modular, open source suite
for stellar evolution



Bill Paxton, father of MESA

- MESA stellar evolution code: mesa.sourceforge.net
- MESA instrument papers ([Paxton et al. 2011](#), [2013](#), [2015](#))

MESA



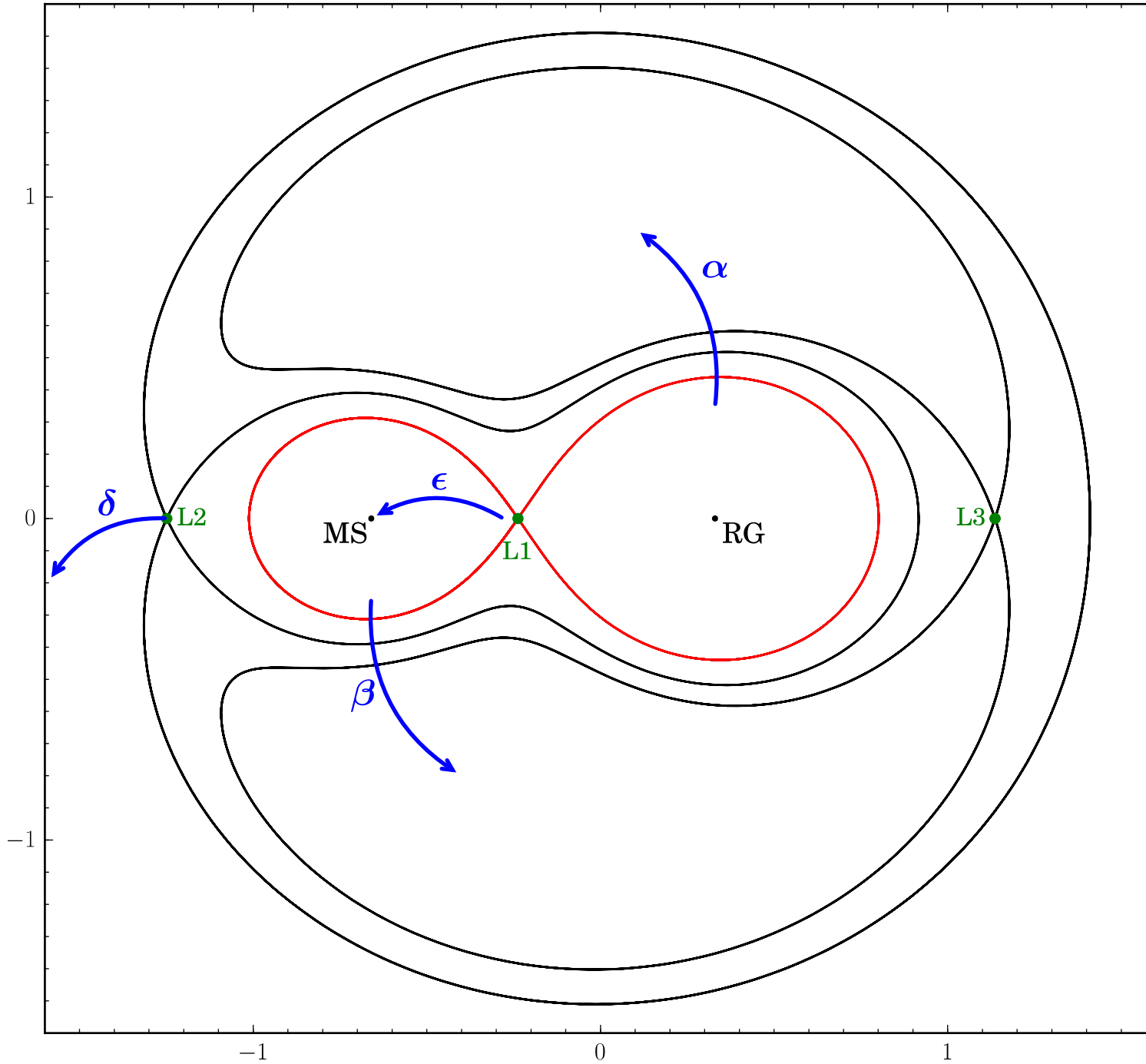
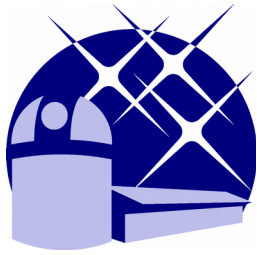
Modules for Experiments in Stellar Astrophysics

Binary Module

- 2 stars evolved at the same time
- Implicit mass transfer (Ritter & Kolb)
- Tides (Zahn 1977)
- Circularisation (Hut 1981, Zahn 1988)
- Magnetic braking
- Gravitational waves
- Angular momentum accretion

Interaction mechanisms presented here will be added in the future versions of MESA

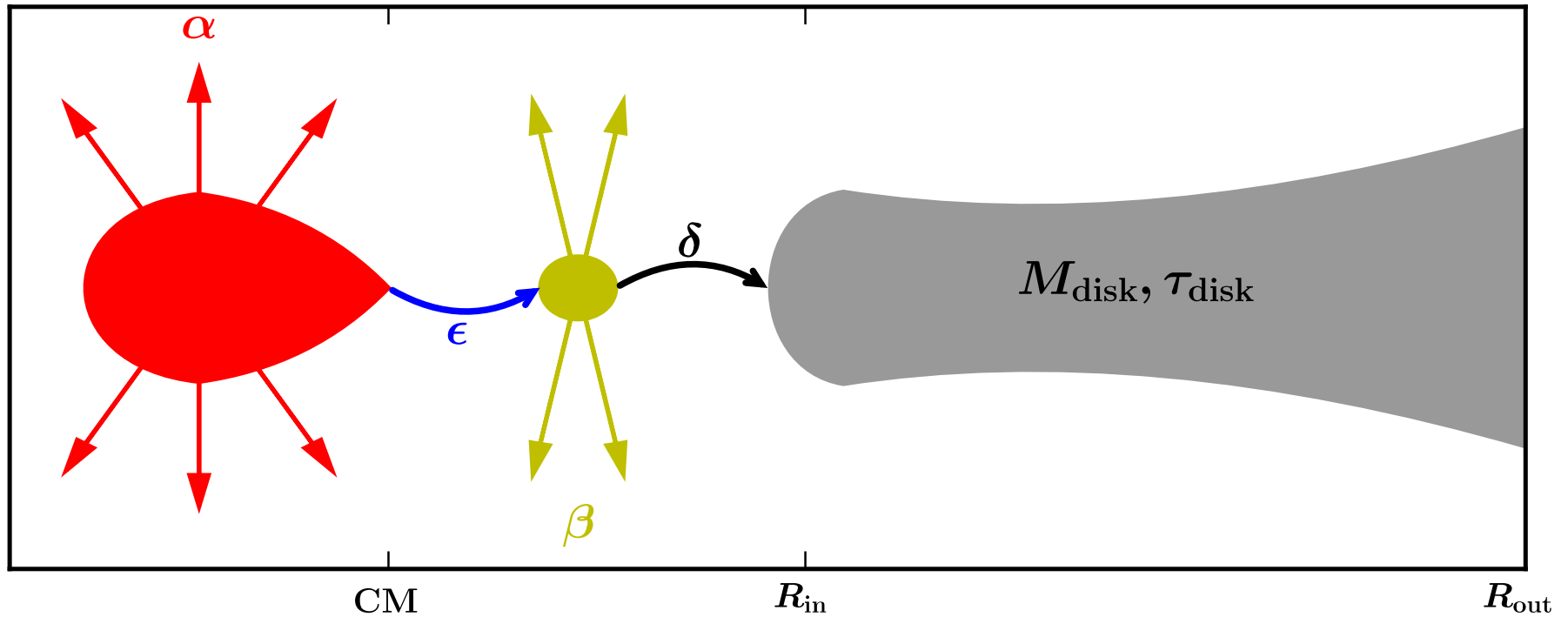
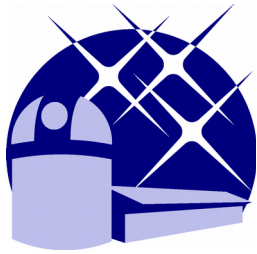
RLOF mass-loss fractions



Mass loss fraction
from Tauris et al.
2006

α : jeans mode
 β : isotropic
re-emission
 δ : circumbinary
coplanar toroid
 ϵ : accretion

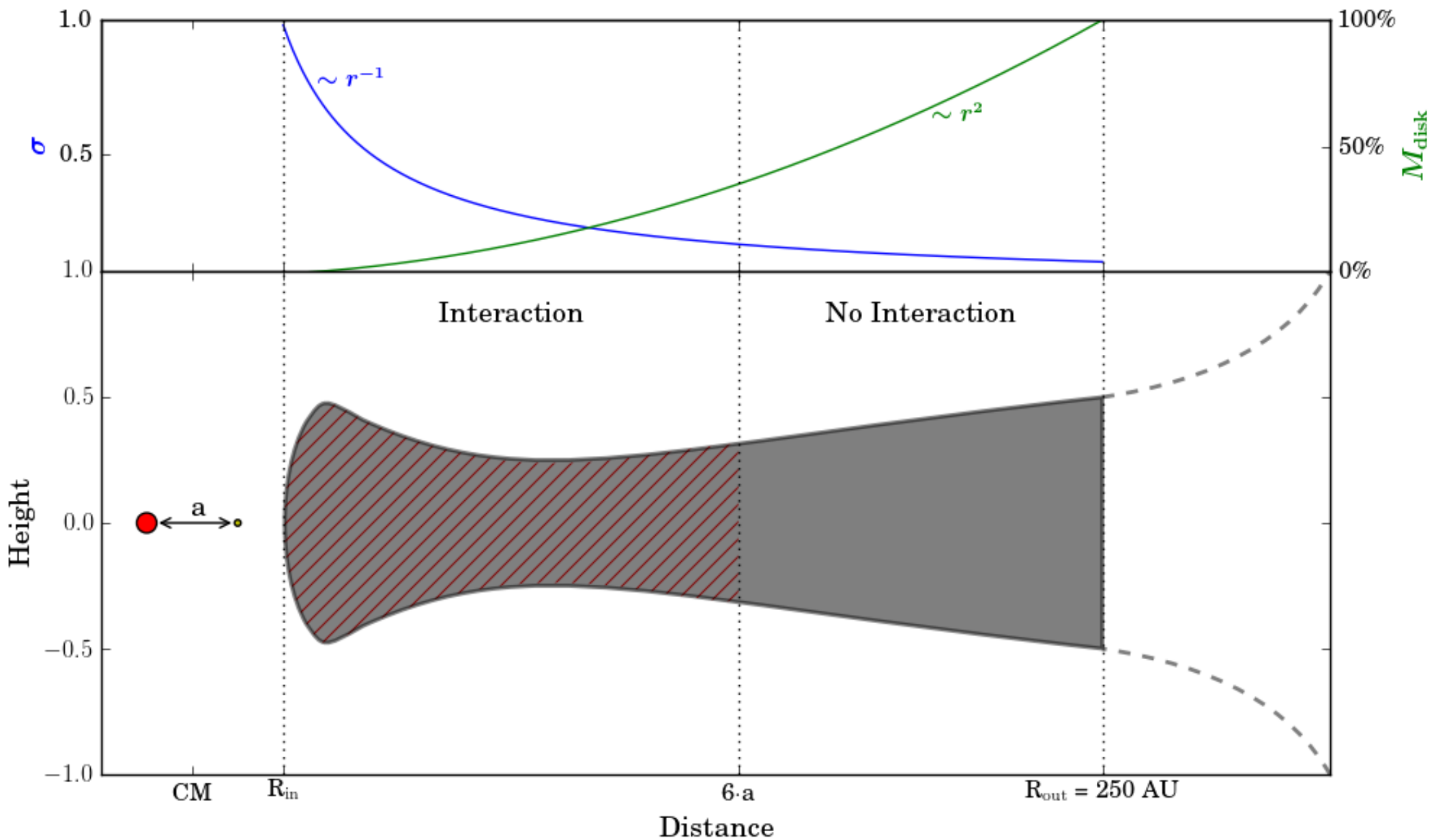
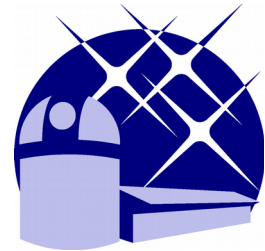
RLOF mass-loss fractions



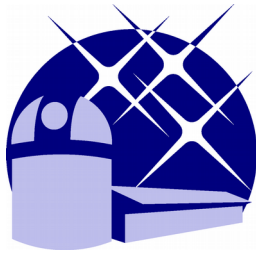
Mass takes away the angular momentum from where it is lost:
donor, companion or L2

δ fraction feeds the CB disk

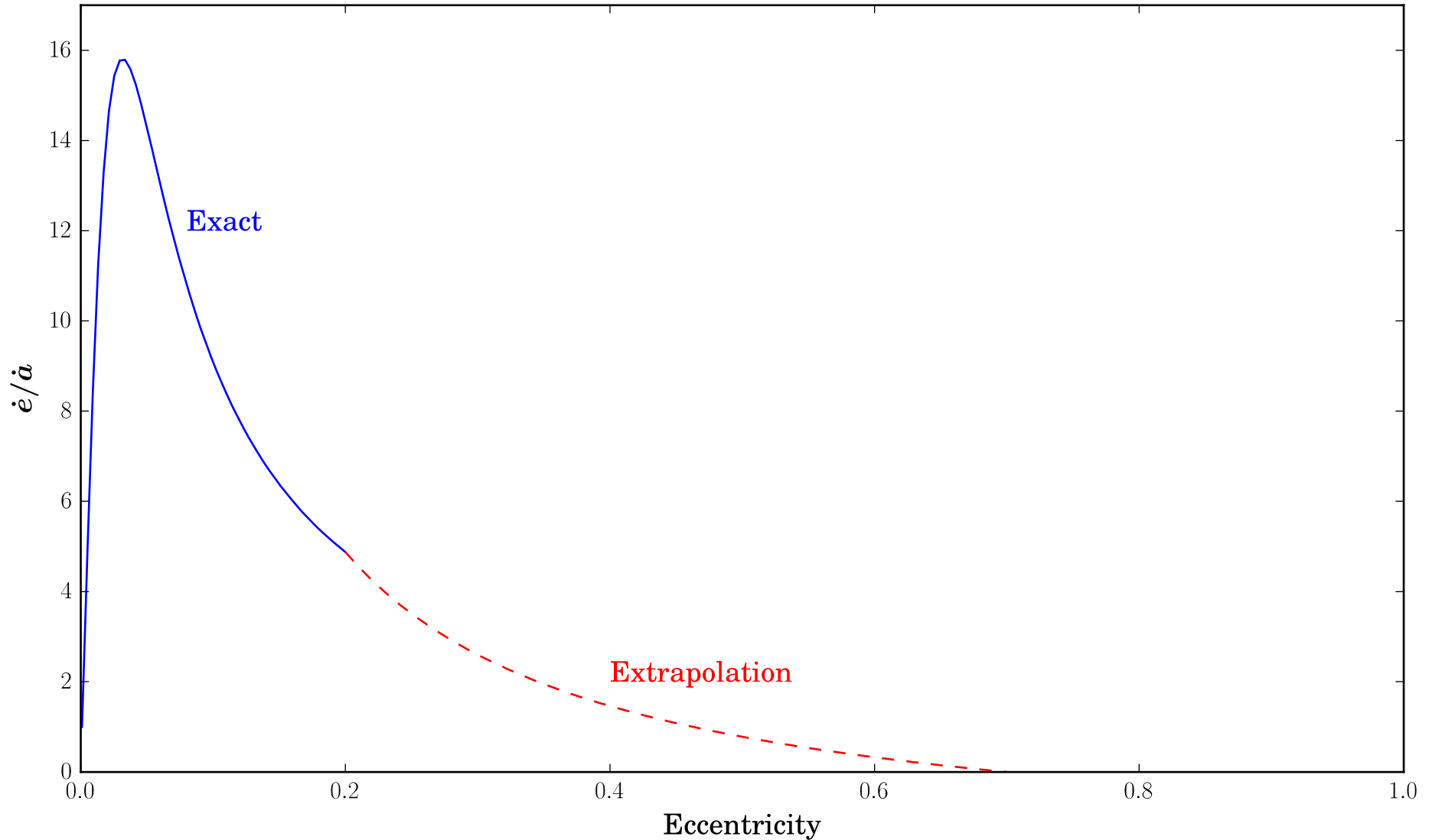
Circumbinary Disk parameters



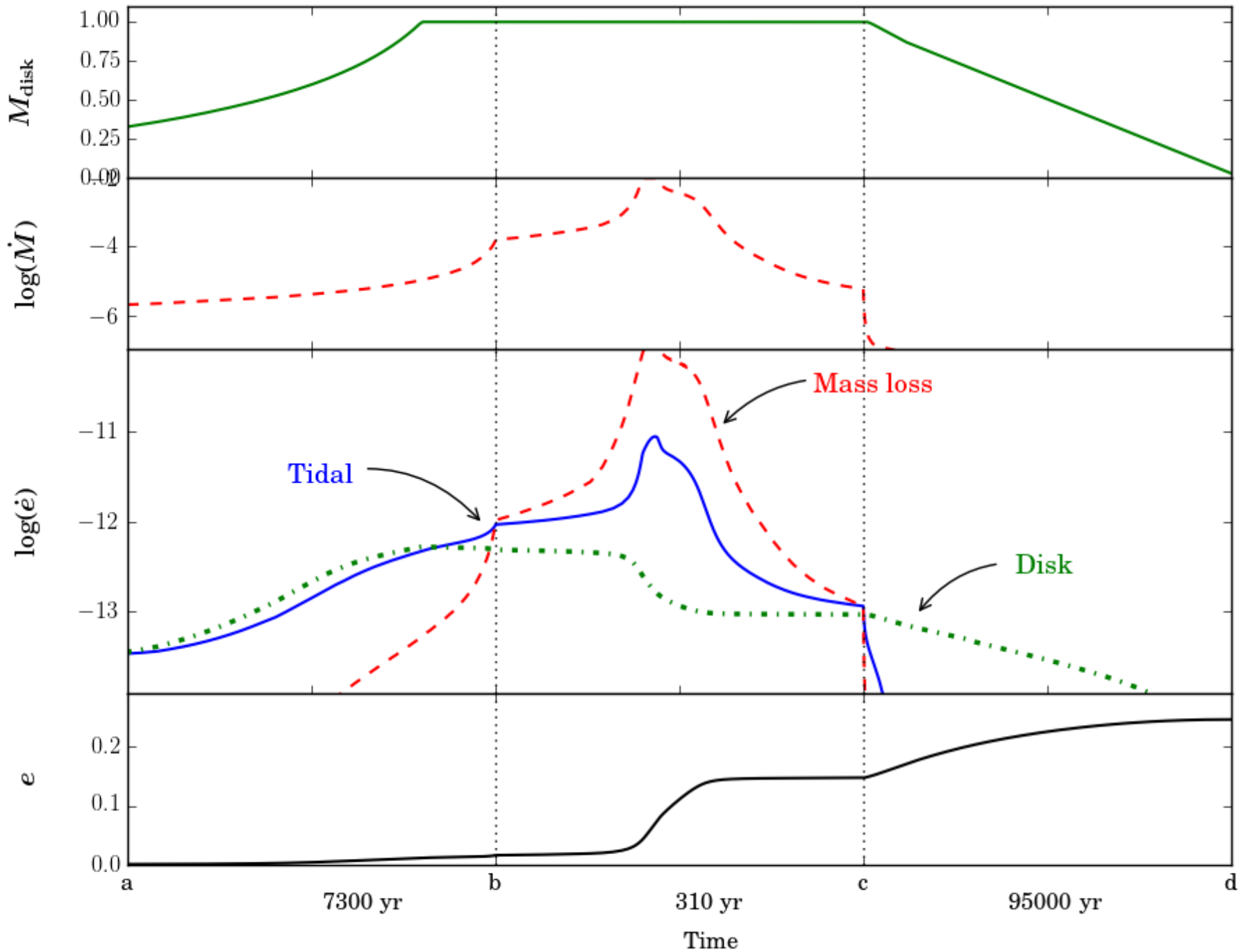
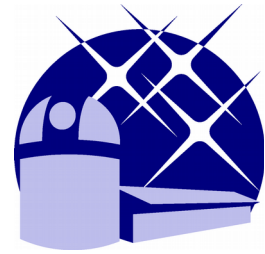
Circumbinary Disk edot



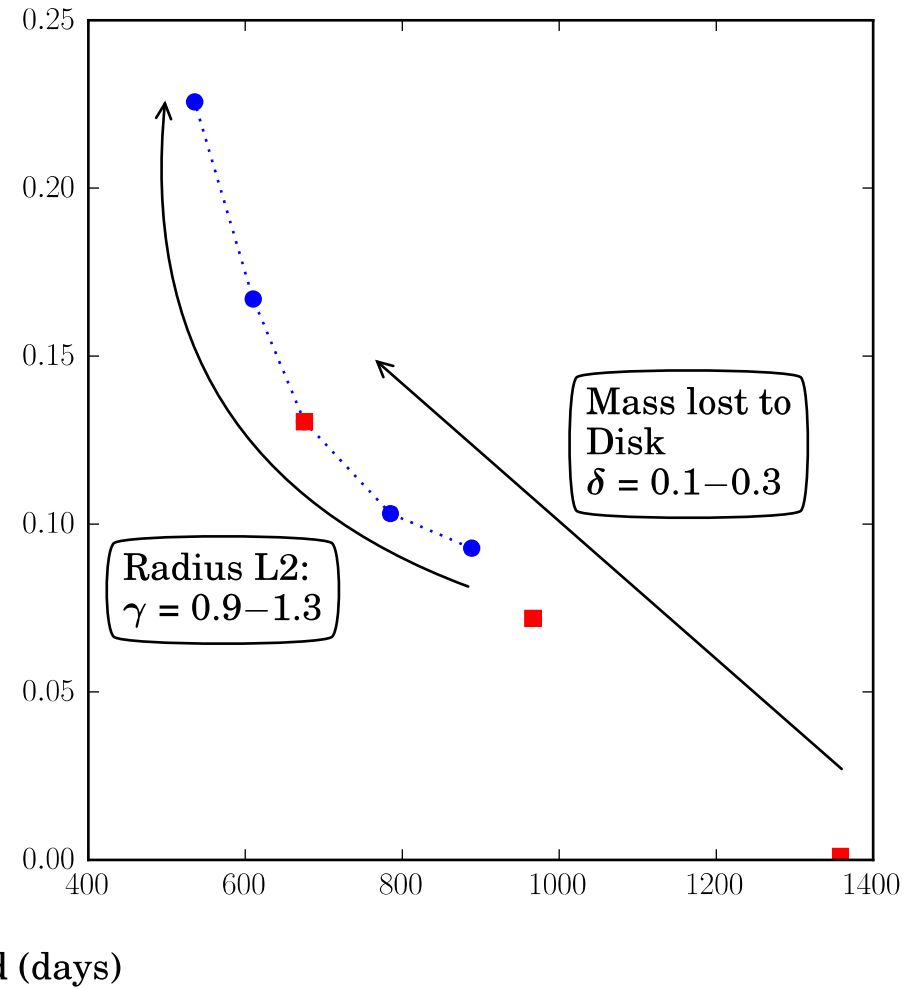
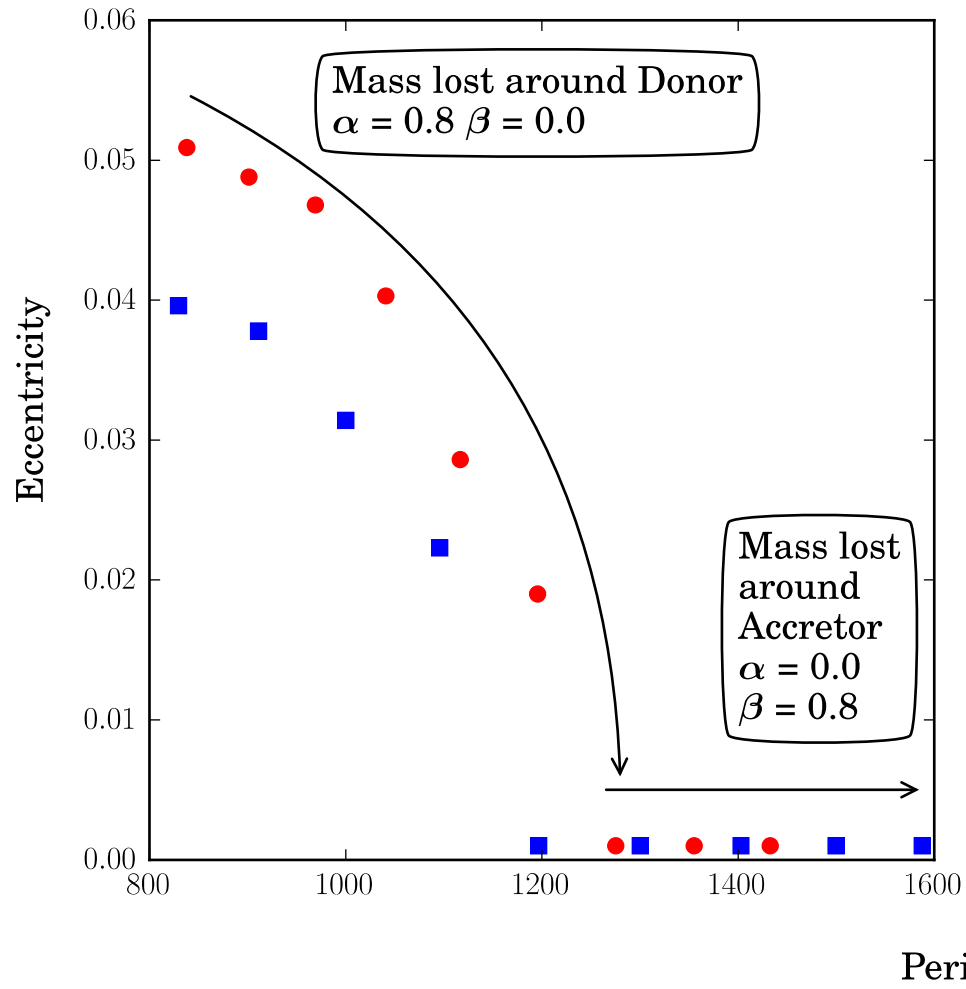
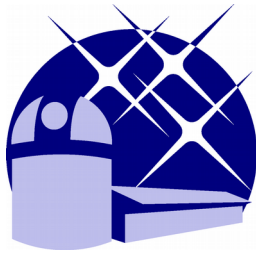
Eccentricity pumping depends on the orbital eccentricity



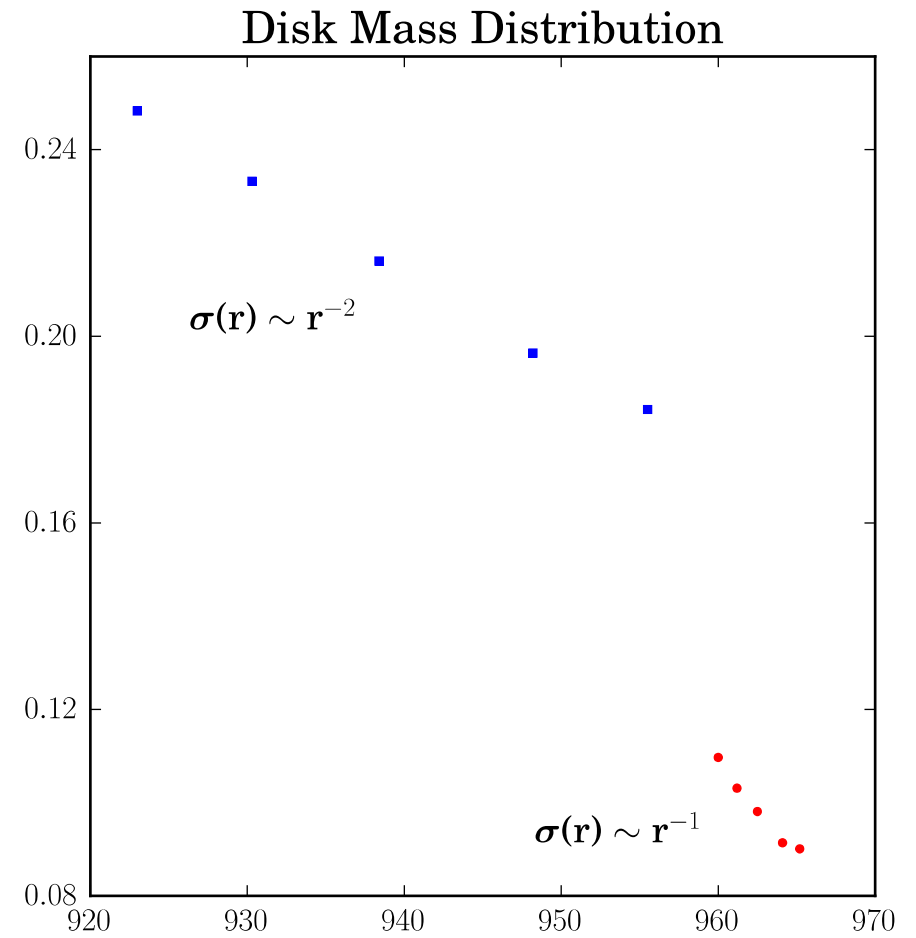
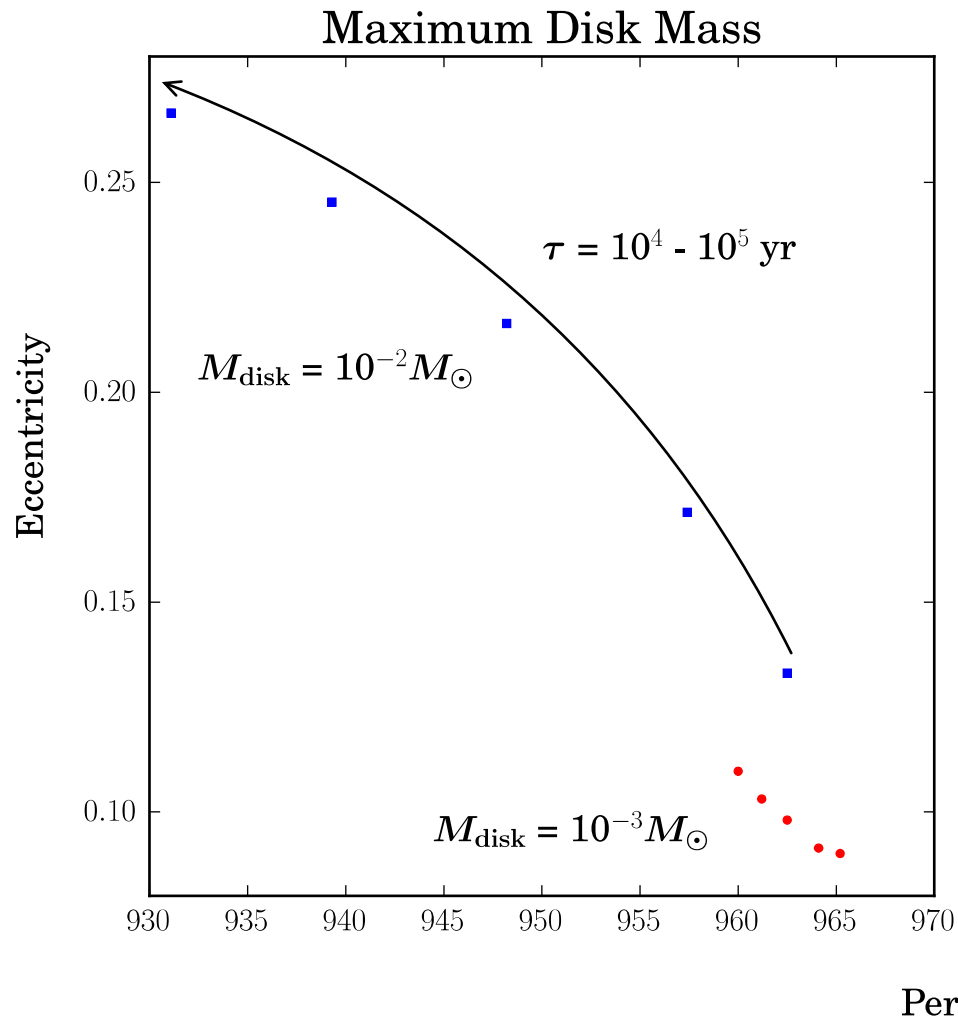
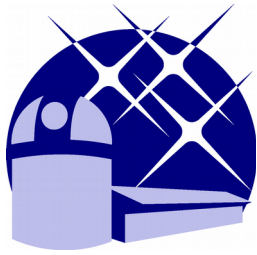
Eccentricity evolution in CB disk model



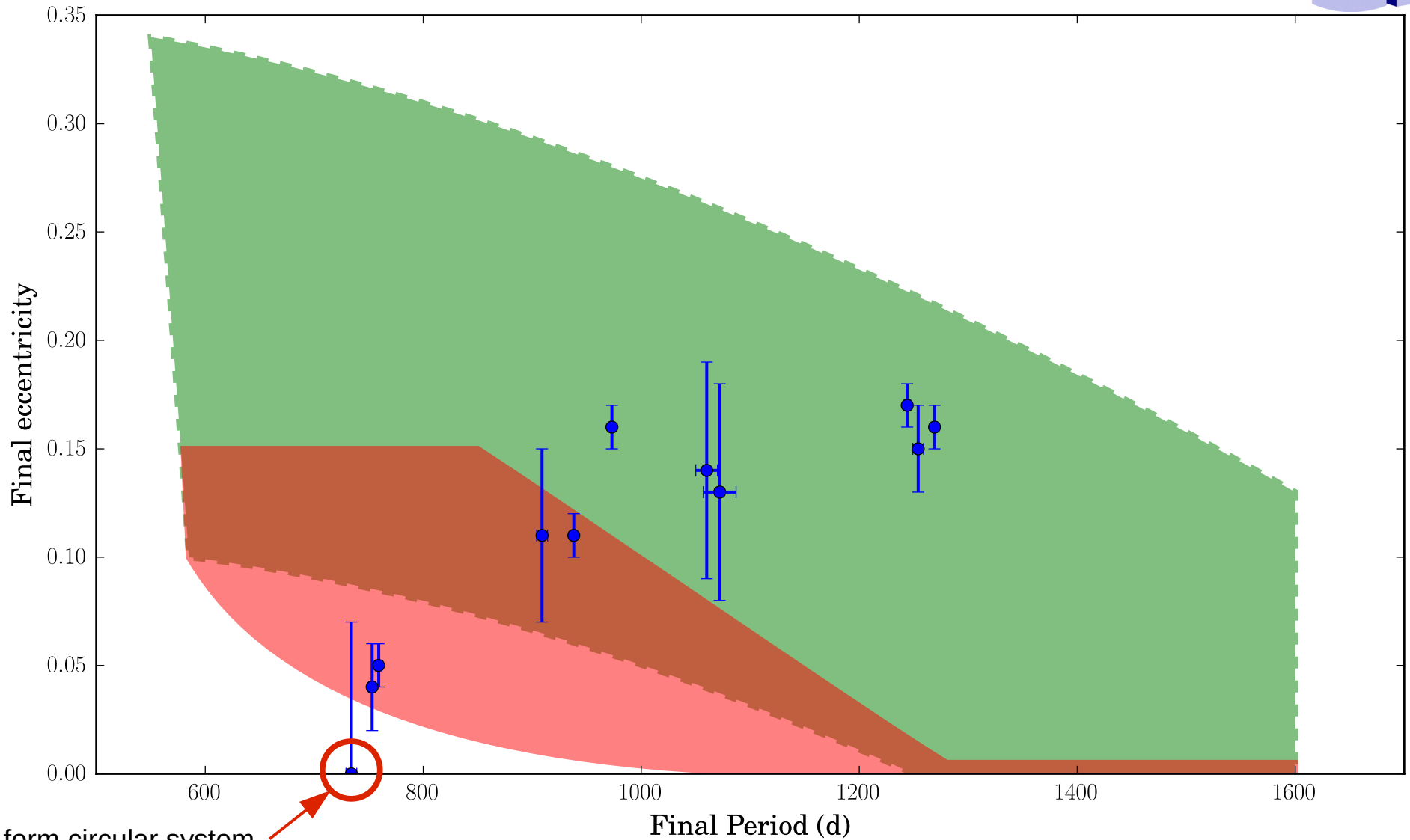
Parameter Effect - RLOF



Parameter Effect - CB disk



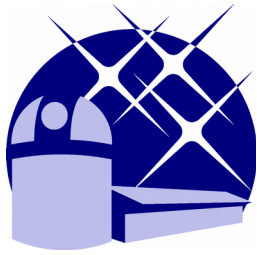
Period - Eccentricity Disk & RLOF



1) Only models with phase dependent RLOF

2) Models with phase dependent RLOF and a CB disk

Results & Conclusions



Created a small test sample for binary interaction mechanisms

MESA

Models allow for observed systems, but don't predict them.

Future prospects:

- Connection to He-WD and dust post-RGB binaries
- Population synthesis studies
- Search for evidence for CB disks
- Continue observing wide sdB binaries

This work was published as: Vos et al. 2015, A&A, 579A, 49V
<http://arxiv.org/abs/1505.03293>