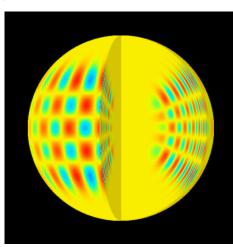
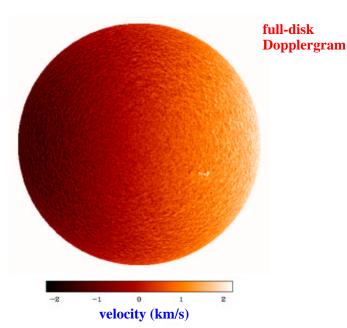
HELIOSEISMOLOGY (I)

acoustic mode in the Sun (p mode n=14, 1 -20)





STRUCTURE OF THE SUN

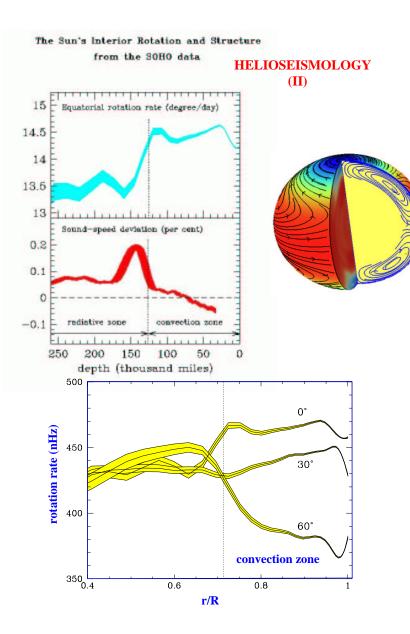
- The Sun is the only star for which we can measure internal properties \rightarrow test of stellar structure theory
- Composition (heavy elements) from meteorites
- Density, internal rotation from helioseismology
- Central conditions from neutrinos

HELIOSEISMOLOGY

- The Sun acts as a *resonant cavity, oscillating* in millions of (acoustic, gravity) modes (like a bell)
- \rightarrow can be used to reconstruct the internal density structure (like earthquakes on Earth)
- oscillation modes are *excited* by *convective eddies*
- periods of typical modes: 1.5 min to 20 min
- velocity amplitudes: $\sim 0.1 \, m/s$
- need to measure *Doppler shifts* in spectral lines relative to their width to an accuracy of $1:10^6$
 - > possible with good spectrometers and long integration times (to average out noise)

Results

- density structure, sound speed
- \bullet depth of outer convective zone: $\sim 0.28\,R_\odot$
- *rotation* in the core is *slow* (almost like a solid-body)
 → the core must have been spun-down with the envelope (efficient core–envelope coupling)



SOLAR NEUTRINOS

- Neutrinos, generated in solar core, escape from the Sun and carry away 2-6% of the energy released in H-burning reactions
- they can be observed in underground experiments
 → direct probe of the solar core
- neutrino-emitting reactions (in the pp chains)

 $\begin{array}{rll} {}^{1}H + {}^{1}H & \to \ {}^{2}D + e^{+} + \nu & E^{max}_{\nu} = 0.42 \, Mev \\ {}^{7}Be + e^{-} & \to \ {}^{7}Li + \nu & E^{max}_{\nu} = 0.86 \, Mev \\ {}^{8}B & \to \ {}^{8}Be + e^{+} + \nu & E^{max}_{\nu} = 14.0 \, Mev \end{array}$

• The *Davis experiment* (starting around 1970) has shown that the neutrino flux is about a factor of 3 lower than predicted \rightarrow the solar neutrino problem

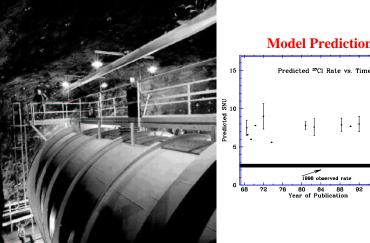
The Homestake experiment (Davis)

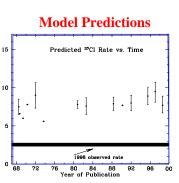
- *neutrino detector:* underground tank filled with 600 tons of Chlorine ($C_2 Cl_4$: dry-cleaning fluid)
- some neutrinos react with Cl

 $u_{\mathrm{e}} + ~^{37}\!\mathrm{Cl}
ightarrow ~^{37}\!\mathrm{Ar} + \mathrm{e^-} - 0.81\,\mathrm{Mev}$

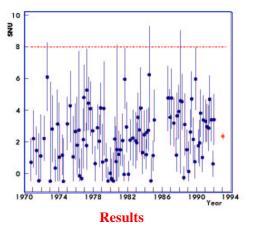
- \bullet rate of absorption $\sim 3 \times 10^{-35} \, {\rm s}^{-1}$ per $^{37}\!Cl$ atom
- every 2 months each ³⁷Ar atom is filtered out of the tank (expected number: 54; observed number: 17)
- \bullet caveats
 - b difficult experiment, only a tiny number of the neutrinos can be detected
 - \triangleright the experiment is only sensitive to the most energetic neutrinos in the ⁸B reaction (only minor reaction in the Sun)

The Davis Neutrino Experiment



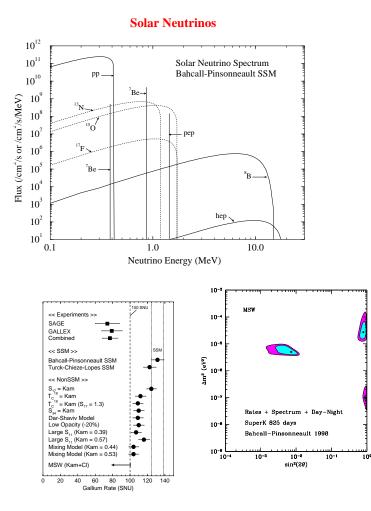


Homestake Mine (with Cl tank)



Proposed Solutions to the Solar Neutrino Problem

- dozens of solutions have been proposed
- 1) Astrophysical solutions
 - ▷ require a *reduction in central temperature* of about 5% (standard model: 15.6×10^6 K)
 - \triangleright can be achieved if the solar core is mixed (due to convection, rotational mixing, etc.)
 - \triangleright if there are no nuclear reactions in the centre (inert core: e.g. central black hole, iron core, degenerate core)
 - \triangleright if there are *additional energy transport* mechanisms (e.g. by WIMPS = weakly interacting particles)
 - ▷ most of these astrophysical solutions also change the density structure in the Sun \rightarrow can now be ruled out by helioseismology
- 2) Nuclear physics
 - \triangleright errors in *nuclear cross sections* (cross sections) sometimes need to be revised by factors up to $\sim 100)$
 - \triangleright improved experiments have *confirmed the nuclear* cross sections for the key nuclear reactions



- 3) Particle physics
 - \triangleright all neutrinos generated in the Sun are *electron neutrinos*
 - \triangleright if neutrinos have a *small mass* (actually mass differences), neutrinos may change type on their path between the centre of the Sun and Earth: *neutrino oscillations*, i.e. change from electron neutrino to μ or τ neutrinos, and then cannot be detected by the Davis experiment
 - ▷ vacuum oscillations: occur in vacuum
 - b matter oscillations (MSW [Mikheyev-Smirnov--Wolfenstein] effect): occur only in matter (i.e. as neutrinos pass through the Sun)

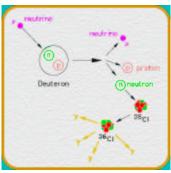
RECENT EXPERIMENTS

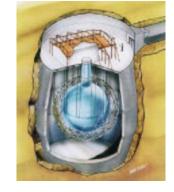
- resolution of the neutrino puzzle requires more sensitive detectors that can also detect neutrinos from the main pp-reaction
- 1) The Kamiokande experiment (also super-Kamiokande)
 - b uses 3000 tons of ultra-pure water (680 tons active medium) for

 $u + e^- \rightarrow \nu + e^- \text{ (inelastic scattering)}$

- \triangleright about six times more likely for $\nu_{\rm e}$ than ν_{μ} and ν_{τ}
- ▷ *observed flux:* half the predicted flux (energy dependence of neutrino interactions?)

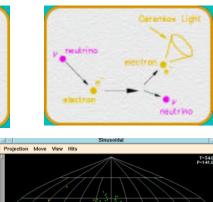
The Sudbury Neutrino Observatory

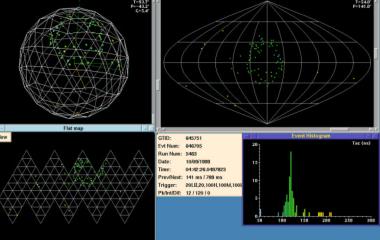




1000 tons of heavy water







- 2) The Gallium experiments (GALLEX, SAGE)
 - b uses Gallium to measure low-energy pp neutrinos directly
 - $u_{
 m e} + ^{71} {
 m Ga}
 ightarrow ^{71} {
 m Ge} + {
 m e^-} 0.23 \, {
 m Mev}$
 - $ightarrow results: ext{ about 80 \pm 10 SNU vs. predicted } 132 \pm 7 \ ext{SNU (1 SNU: 10^{-36} interactions per target atom/s)}$
- 3) The Sudbury Neutrino Observatory (SNO)
 - \triangleright located in a deep mine (2070 m underground)
 - \triangleright 1000 tons of pure, *heavy water* (D₂O)
 - \triangleright in a crylic plastic vessel with 9456 light sensors/photo-multiplier tubes
 - b detect Cerenkov radiation of electrons and photons from weak interactions and neutrino-electron scattering
 - > results (June 2001): confirmation of neutrino oscillations (MSW effect)?